

R. M. Thorne showed that the whistler-mode plasma wave turbulence observed in Io's torus by Voyager 1 would produce rapid precipitation of torus electrons into Jupiter's atmosphere.



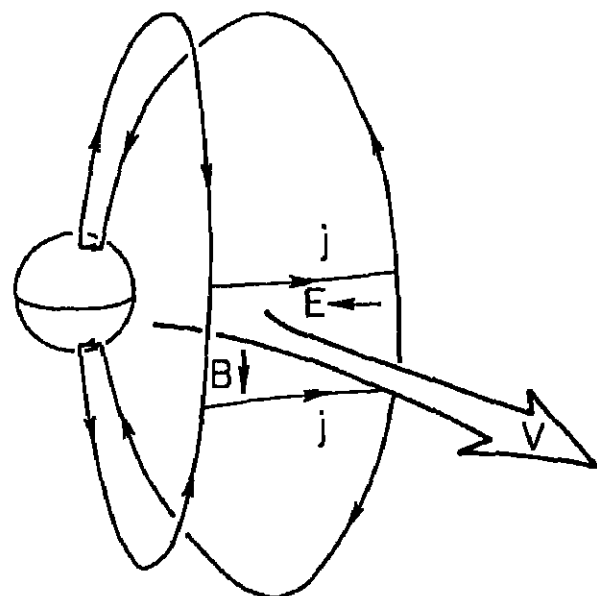


Fig. 3. The longitudinally asymmetric injection of plasma in the Io torus produces a partial ring current, which closes by means of field-aligned currents connecting to Pedersen currents in Jupiter's ionosphere. The closure of the Pedersen current requires an electric field  $E$ , which is associated with outward drift  $v = E/B^2$ . An inward return flow toward Jupiter occurs in the opposite longitude sector (not shown). The convection pattern corotates with Jupiter and provides a mechanism for rapid outward transport of the Io torus plasma (T. W. Hill), while the associated field-aligned currents may produce decametric or kilometric radiation (A. J. Dessler).

sphere. In addition to accounting for the observed auroral emissions from Jupiter, this precipitation would imply the existence of a rapid radial transport or local acceleration mechanism to replace the electrons lost into the atmosphere on a time scale of 1 day.

T. V. Johnson discussed a model of electrostatic charging (by electron impact) of dust particles carried aloft by Io's eruptive plumes. He estimates that submicrometer-size dust particles may be charged up to  $\sim 10$  V, sufficient for the Lorentz force to overcome Io's gravity and allow the dust particles to escape directly to the magnetosphere. The subsequent breakup of such particles would directly provide a distributed source of sulfur and oxygen to the torus. The mass output of the plumes ( $\sim 1000$  kg/s) appears sufficient to maintain the observed plasma torus against loss by outward transport.

L. J. Lanzerotti reported laboratory measurements of the sputtering of  $SO_2$  ice by incident 1.5-MeV ions, indicating that incident  $O^+$  ions produce a remarkably high yield of  $\sim 4000$  sputtered atoms per incident ion. This result suggests that the Io torus may sustain itself to a large degree by heavy-ion sputtering of material from the  $SO_2$ -enriched surface of Io.

#### Low-Frequency Radio Emissions

'Decametric' (wavelength  $\lambda \sim 10$  m) radio emissions from Jupiter have been observed extensively by earthbound radio telescopes for over two decades, and they have provided

much indirect evidence about Jupiter's magnetosphere, and especially its interaction with Io. The occurrence rate and intensity of decametric radio storms depend strongly on both the Jovian longitude of the observer and the orbital phase of Io relative to the observer. These correlations have given rise to models wherein the radiation is emitted in a narrow conical beam nearly perpendicular to the magnetic field near the foot of the Jovian magnetic flux tube that intersects Io. The energy source for the emission is widely considered to be an electron beam associated with the magnetic-field-aligned current that connects Io with Jupiter's ionosphere, the current being driven by the EMF associated with the relative motion between Io and the corotating magnetospheric plasma. Radio receivers on Earth-orbiting satellites, and on Voyagers 1 and 2, have now extended these decametric observations to longer wavelength regions of the spectrum—heliometric ( $\lambda \sim 100$  m) and kilometric ( $\lambda \sim 1$  km). The Voyager receivers have also provided a new viewing perspective on the post-encounter trajectories in the predawn sector of Jovian local time.

T. D. Carr reviewed recent ground-based and Voyager-based radio observations, noting that the Io-orbital phase control tends to disappear at lower frequencies (heliometric and kilometric) and that new Voyager observations indicate an apparent local-time dependence superimposed on the known Jovian-longitude dependence of the radio sources (see below). He also pointed out that the location(s) of the source regions has not been definitively established for any of the low-frequency radio emissions. It was noted by J. Alexander that the observed frequency and polarization characteristics would place most if not all known emission sources at low altitudes in Jupiter's northern hemisphere (with the possible exception of the narrow-band kilometric component; see below).

J. R. Thieman observed that the decametric (10 MHz) Io-dependent sources appeared at the same Jovian longitudes in pre- and post-encounter Voyager observations as they did in Earth-based observations, thus supporting the traditional view that the radio sources corotate with Jupiter and are independent of local time. On the other hand, he noted that the Io-independent component seems to shift toward smaller longitudes in the post-encounter observations, as if there were a local-time dependence of the strength of the various source regions. This result would present theoretical difficulties comparable to those noted above in connection with the apparent local-time dependence of the UV brightness of the Io torus.

Similarly, the broadband kilometric component, discovered by Voyager 1 and described by M. D. Desch, exhibits an apparent local-time dependence. The pre- and post-encounter observations are clearly different with regard to polarization sense (LH before encounter, RH after) as well as emission probability and intensity (smaller after encounter than before). These results were interpreted in terms of a model in which the post-encounter observations are attributed to over-the-pole viewing of the same dayside source that produces the pre-encounter observations when viewed directly from local noon (Figure 4). Some skepticism was expressed about the plausibility of this over-the-pole viewing geometry, but an alternative explanation of the difference between pre- and post-encounter observations in terms of a latitude effect also has implausible features. A satisfactory explanation for the observations was not immediately apparent.

M. L. Kaiser described a narrow-band kilometric wave component (also discovered by Voyager 1) whose source apparently rotates 3%–5% slower than the rigid System III (1965) rotation rate that is characteristic of all other known Jovian radio sources. From this rotation lag, and from considerations of propagation and viewing geometry, Kaiser infers that the source of the narrow-band kilometric component is located near the outer edge of the Io plasma torus in the magnetospheric equatorial plane rather than at the high-latitude Jovian topside ionosphere, as is thought to be the case for the other low-frequency sources.

It was noted by Alexander and by C. K. Goertz that the decametric emissions, and especially the Io-independent component, often occur at frequencies above the cyclotron frequency of electrons in the strongest magnetic field accessible to trapped electrons, the implication being that the emission is caused not by trapped electrons but by precipitating electrons.

A. J. Dessler noted that the magnetic anomaly model, in which plasma is produced preferentially in a particular active sector of Jovian longitude, predicts a partial ring current in the Io torus (Figure 3), which in turn produces magnetic-field-aligned current densities of  $\sim 1/4$  A/km<sup>2</sup>, similar to the magnitude of current density that is known to produce terrestrial kilometric radiation.

M. L. Goldstein proposed a mechanism to account for the arc-shaped features in the dynamic spectra observed by the Voyagers, involving an emission cone angle that depends on frequency. D. A. Gurnett further proposed that the observed multiplicity of these arc emission features might be produced

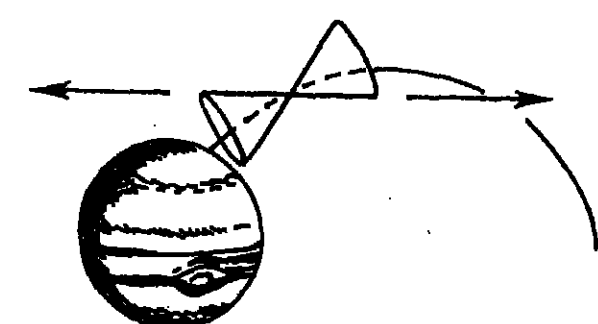


Fig. 4. Beaming model proposed by M. D. Desch in which broad-band kilometric radiation is emitted along the surface of a cone that may intersect the observer either on the same side of Jupiter (to the right) or over Jupiter's pole (to the left). It is proposed that the emission is most probable and/or most intense on the dayside of Jupiter; the difference between direct and over-the-pole viewing geometries would then explain the differences between pre-encounter and post-encounter Voyager observations.

by multiple bounces, between northern and southern Jovian hemispheres, of a large-amplitude Alfvén wave produced by the interaction between Io and the corotational magnetospheric flow.

Kaiser also announced the first definitive observation (by the Voyager 1 and 2 planetary radio astronomy experiments) of radio emissions from Saturn. The emissions were observed at 200 kHz, and Kaiser noted that Saturn as a radio emitter looks much more like Earth than like Jupiter.

#### Plasma Flow and Rotational Dynamics

The Pioneer 10 and 11 encounters (1973–1974) confirmed earlier suspicions that the corotation of the magnetospheric plasma with the planet would produce more important dynamical effects at Jupiter than at Earth, primarily because of the larger size and faster rotation rate of Jupiter. Voyager 1 and 2 (1979) measurements, on the other hand, have now confirmed recent theoretical suggestions that corotation in Jupiter's magnetosphere is imperfect, with the angular velocity decreasing with increasing distance from Jupiter, owing to the weakness of the atmosphere-magnetosphere coupling and the rapid injection of plasma from Io into the magnetosphere. Nevertheless, corotation has important effects on the magnetosphere, including the 10-hour spin modulation of energetic particles and magnetic fields in the magnetosphere and the spin-periodic ejection of energetic particles into interplanetary space. The centrifugal force of corotation is responsible, at least in part, for inflating the outer magnetosphere into a disklike field geometry, and several authors have suggested that the corotational centrifugal force causes the magnetic field to open at some distance to allow the escape of accumulated plasma in the form of a 'planetary wind'. Pioneers 10 and 11 had discovered persistent modulations of the magnetic field and particle fluxes at Jupiter's 10-hour rotation period, both inside and outside the magnetosphere, as illustrated schematically in the cover figure A. The 'magnetic disc model' attributes these 10-hour variations within the magnetosphere to the diurnal precession of Jupiter's magnetic axis about its spin axis, which causes a periodic wobble of the disc field configuration (cover figure B). The 'magnetic anomaly model' attributes the 10-hour variations both inside and outside the magnetosphere to the corotation of a longitudinal plasma asymmetry (cover figure C) that causes the planetary wind and related processes to vary with the spin period. The Pioneer trajectories lay almost entirely outside the range of latitudinal wobbling of the proposed disc (cover figure B), and the disc model would thus predict only one maximum and one minimum per 10-hour cycle, as would the anomaly model. The two models were thus indistinguishable on the basis of Pioneer data. The trajectories of both Voyager spacecraft lay within the latitude range of the hypothetical rigidly wobbling disc (cover figure D), where in the disc model would predict two maxima and two minima per rotation (cover figure E, solid line), while the anomaly model would still predict one (cover figure E, dashed line). Voyager data within 80  $R_J$  are generally consistent with the disc model signature, while a number of observations (some of them noted below) appear to require the existence of a corotating anomaly as well.

V. M. Vasyliunas compared Voyager 1 and 2 plasma measurements obtained by the plasma science experiment (PLS) and by the low-energy charged-particle experiment (LECP). The departures from rigid corotation reported by the PLS are consistent with a predicted corotation lag caused by the inertial drag of plasma continuously injected by Io. The LECP results reported by S. M. Krimigis et al., on the other hand, indicate strict corotation to greater distances than would seem to be consistent with the PLS results, and there remains some controversy as to the degree to which corotation is enforced in the outer magnetosphere (see section entitled 'Outstanding Issues and Controversies,' below). R. L. Mott reported from PLS measurements that departures from corotation become significant beyond the orbit of Io at 6  $R_J$ , and that the PLS results would be consistent with the 5% lag in the outer torus inferred by Kaiser from the narrow-band kilometric observations.

Vasyliunas also noted that the antisunward flow reported by the LECP experimenters in the distant magnetospheric tail was consistent with theoretical expectations based on the concept of a planetary wind (Figure 5). Apparently this wind consists primarily of heavy ions from the Io torus, although the composition of the wind was not measured directly.

R. P. Lepping reported Voyager observations of spin-periodic perturbations of the magnetic field in the magnetosheath (just outside the magnetosphere). These observations appear to support the basic premise of the magnetic anomaly model. A. W. Schardt presented an analysis of Voyager crossings of the magnetospheric current sheet (disc). His analysis indicated a reduced Alfvén wave propagation speed, and hence an enhanced plasma mass density, in the active sector as defined by the magnetic anomaly model. I. de Pater showed an analysis of high-frequency radio observations that suggests the existence of a rotating 'hot spot'—a localized excess concentration of relativistic synchrotron-radiating electrons near the planet—in roughly the same active sector as that defined by the magnetic anomaly model.

#### Particle Acceleration, Diffusion, and Loss

The classical mechanism for particle acceleration in planetary magnetospheres is the 'betatron' process whereby charged particles are transported inward into regions of higher magnetic field strength. If the transport occurs adiabatically with respect to the first and second adiabatic invariants, the particle's cyclotron energy increases in proportion to the increase of equatorial magnetic field strength. The process is then analogous to the adiabatic compression of an ideal gas. The inward transport may occur through randomly phased, inward and outward motions of magnetic flux tubes ('radial diffusion') or through a systematic circulation ('convection')

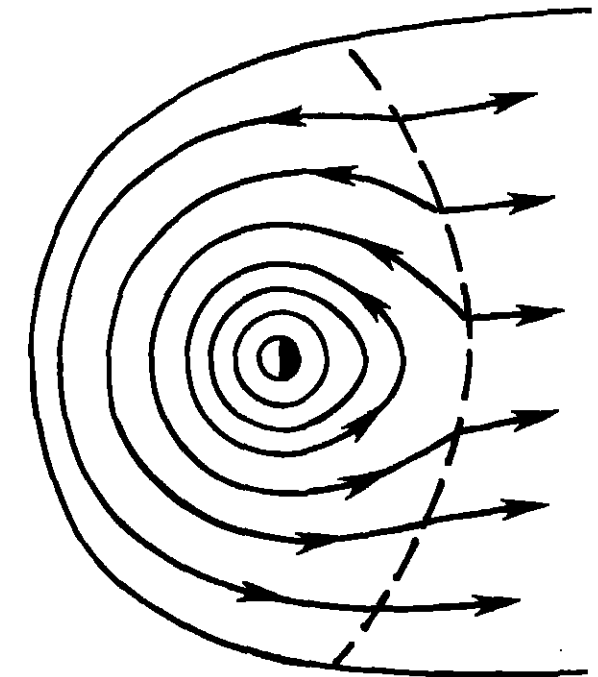


Fig. 5. Schematic illustration of the generation of a planetary wind on Jupiter's nightside (equatorial view, with the sun to the left). As illustrated here, plasma rotates on closed magnetic flux tubes throughout the dayside magnetosphere, but the outermost flux tubes are forced open to allow plasma escape in the dusk sector of the tail. Reconnection of these open flux tubes in the dawn sector of the tail produces the apparent divergence of the flow lines away from the reconnection region (dashed line), resulting in primarily rotational flow within the dashed line and tailward flow beyond the dashed line. As pointed out by V. M. Vasyliunas, this theoretical scenario is consistent with observations reported by the Voyager low-energy charged-particle experiments.

which particles are transported inward on one half of the flow cycle and outward on the other half. Pioneer 10 and 11 observations had confirmed the importance of betatron acceleration through radial diffusion in Jupiter's magnetosphere, as well as the expected role of Io in accelerating particles to energies of several hundred kiloelectron volts. The Pioneer observations also, however, demonstrated the need for an additional, faster, and more potent mechanism to produce the several million electron volt particles observed in the outer magnetosphere (where the magnetic field is relatively weak), as well as the similar particles that are ejected into interplanetary space.

T. W. Hill reviewed recent theoretical mechanisms that focus on the problem of explaining the energetic particles (several million electron volts) observed in the outermost magnetosphere and in interplanetary space. One is a reconnection mechanism wherein particles are recycled through the betatron process by alternating cycles of adiabatic compression and nonadiabatic (rapid) expansion. The 'magnetic pumping' process also involves alternating cycles of adiabatic acceleration and nonadiabatic relaxation, but it utilizes the day-night variation of magnetic field strength along a corotational drift path rather than the radial gradient of field strength. Both of these processes are capable of producing million electron volt energies in the outer magnetosphere, but both are intrinsically slow processes that apparently cannot sustain the high rate of ejection of energetic particles into interplanetary space and the rapid refilling of the outer magnetosphere.

A potentially powerful and rapid acceleration mechanism involves magnetic merging or field annihilation in the current sheet of the outer magnetosphere. This process provides, in principle, a means of rapidly tapping the energy stored in the highly stressed magnetic field configuration. The merging theory is not sufficiently developed to make quantitative predictions of the expected particle energy spectrum, but observations in Earth's magnetospheric tail suggest that the merging process is capable of producing the required high energies on the required short time scale. R. W. Fillius reviewed an important analysis of Pioneer observations by A. W. Schardt and colleagues, showing evidence of local acceleration of energetic particles in the outer-magnetospheric current sheet. The observed flow of energetic ions away from the current sheet apparently implies a total acceleration rate of the order of  $10^{14}$  W and a relatively short particle lifetime of the order of the 10-hour Jovian rotation period. These observations can be taken as evidence for a current-sheet field annihilation process.

## News

### R & D Funding in 1981

Although it is widely accepted that federal budgets will be cut during the next year, it is apparent now that R & D support will remain strong for FY 1981. The National Science Foundation's overall budget has been increased by 8% for FY 1981. NASA R & D has gained 7% on its budget. The Department of Commerce R & D budget has been increased by 6%. The huge increases are among the military and defense agencies, some amounting to as much as 37% (Air Force), with an average DOD R & D increase of 19%. The table lists the increases by agency.

The breakdown of the NSF budget for FY 1981 includes a 15% increase in funding for mathematical and physical sciences; 9% for earth, ocean, atmosphere, and astronomical sciences; 10% for Antarctic programs; and 7% for biological and social sciences. The Ocean Margin Drilling Program will increase in budget by 13%.

The NASA budget and future programs will be affected by some interesting twists in congressional legislation. The over-

A. J. Dessler pointed out the existence of a theoretical upper limit to the rate at which energy can be drawn from the centrifugal potential field of corotation, this limit being proportional to the rate at which plasma is injected by Io into the magnetosphere. Using the largest values inferred for this plasma mass injection rate ( $\sim 10^3$  kg/s), this upper limit is dangerously close to the  $10^{14}$  W that was mentioned earlier in reference to both the torus-associated auroral emission and the current-sheet acceleration process.

Fillius and C. E. McIlwain noted that the possible role of large magnetic-field-aligned electric fields has been generally overlooked in theoretical attempts to explain the MeV electrons observed in the outer magnetosphere, although such fields are widely recognized as important in the acceleration of auroral electrons in the terrestrial magnetosphere. Differential rotation between different portions of a given field line can in principle produce megavolt potential drops along the magnetic field.

A. Hasegawa presented a theory of a 'ballooning instability', wherein the centrifugal force of the plasma that streams along the sharply curved magnetic field lines which thread the current sheet forces the magnetic field to stretch to the breaking point. The instability is analogous in some respects to the classical Rayleigh-Taylor instability, with the centrifugal 'gravity field' directed outward. The instability may be instrumental in releasing the planetary wind outflow and in establishing the highly stressed magnetic field configuration that is a prerequisite for the fast magnetic merging process described earlier.

M. Schulz discussed a synergistic interaction between two radiation modes that affect the loss of relativistic electrons in the innermost magnetosphere—a cyclotron instability, which tends to occur near the magnetic equator, and synchrotron emission, which tends to occur more at the high-latitude excursion of a particle bounce trajectory along a magnetic field line. Each radiation mechanism alters the particle velocity distribution in such a way as to enhance the other. The interplay between the two radiation mechanisms was thus aptly described by Schulz as a 'cyclotron-synchrotron maser'.

### Outstanding Issues and Controversies

The first two points listed below were discussed at length in the final session of the conference; these discussions served to bring the issues into clearer focus but did not really resolve them. The remaining questions were identified or emphasized during the course of the conference as critically important issues to be resolved by future research.

There is an apparent discrepancy between plasma flow measurements reported by the two plasma experiments on each of the Voyager spacecraft. The two experiments cover different ranges of particle energies but should presumably give the same value for the components of plasma flow perpendicular to the magnetic field, and in particular, for the corotational component of flow. As was pointed out by Vasyliunas, there is not yet a clear, direct conflict between the two sets of measurements because there is no overlap between the regions of space within which the two experiments have reported flow measurements. However, the PLS has reported significant (up to 50%) departures from ideal corotation between 10 and 40  $R_J$  distance, the degree of departure increasing with increasing distance, as expected theoretically, while the LECP has reported essentially rigid corotation between 40 and  $\sim 100 R_J$ . It is difficult to reconcile these two results theoretically, and a direct comparison of simultaneous results from the two instruments in the same region of space would be desirable in order to decide whether we have an experimental discrepancy on the one hand or a theoretical dilemma on the other.

It is important to establish whether the source of torus plasma is localized near Io or is distributed widely around Io's orbit. In the former case, the low intensity of UV emission observed from Io's vicinity would be hard to reconcile with the large plasma injection rates ( $\geq 10^3$  kg/s) inferred from the optical observations, the intensity of the Io-associated Jovian aurora, and the observed corotation lag. In the latter case, the interesting problem would be to account theoretically for such a widely distributed source.

How persistent are the reported local-time asymmetries in UV torus emissions and certain low-frequency radio emissions? Can these asymmetries be attributed to a latitude rather than a local-time variation? If the local-time variations are real, what do they tell us about the Io-Jupiter interaction?

Is the rapid outward transport of torus plasma better described in terms of a systematic convection pattern or in terms of stochastic radial diffusion?

- What is the energy budget of the Io torus, and what are the sources and sinks of this energy?
- Is the Io-Jupiter current system better described as a steady state circuit closed by ohmic currents in Jupiter's ionosphere or as a system of Alfvén waves only weakly coupled to Jupiter's ionosphere?
- What is the nature of the interaction that produces the Ganymede wake?
- What are the precise locations of the various low-frequency radio sources? (The answer to this question would narrow the choices of relevant emission mechanisms.)
- What are the relative abundances of hot versus cold plasma in the outer magnetosphere and the relative importance of their dynamical effects ('hot' and 'cold' being defined relative to the local corotation energy)?
- What are the relative roles of the disk model versus the magnetic anomaly model in producing 10-hour variations in the outer magnetosphere and beyond?
- What is the nature of the magnetosheath disturbances that are observed to occur with a 10-hour periodicity?
- What mechanism is primarily responsible for the rapid production of energetic particles in the outer magnetosphere and their ejection into interplanetary space?
- Does the planetary wind form just at, or significantly beyond, the distance at which the rotational energy density becomes comparable to the magnetic energy density? Does the planetary wind exhibit a 10-hour periodicity?
- Are there large-scale regions of magnetic-field-aligned electric fields associated with differential rotation? If so, what is their role in energetic particle acceleration?

This list is probably incomplete, but it serves to illustrate the kinds of questions we are learning to ask in this post-Voyager, pre-Galileo era of Jovian magnetospheric research.

### Acknowledgments

I would like to thank all of the conference participants, and especially the session chairmen (P. A. Cloutier, A. J. Dessler, M. L. Goldstein, D. M. Hunten, E. C. Stone), the overview speakers (R. A. Brown, T. D. Carr, F. W. Fillius, C. K. Goertz, D. A. Gurnett, T. W. Hill, G. L. Siscoe, V. M. Vasyliunas), and the program committee (a subset of the above lists) for their indispensable contributions to the success of the conference (and for their tolerance of this imperfect review). I would also like to extend special thanks to the Space Physics and Astronomy Department and its chairman, Alex Dessler, for their sponsorship and support, and to Umbeline Cantu, Georgia Burgess, Rachid Mesli, Tom Tascione, and Huey-Ching Yeh for their logistical support. The preparation of this review was supported in part by NSF grant ATM78-21767.

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an Academy panel. The implications of this congressional move are uncertain.

The overall requests of the Administration were followed almost 100%. The biggest discrepancies between the Administration's request and Congress' approval were in the budgets of the Defense Department, 3% less than requested, and the Environmental Protection Agency, 6% less than requested; however, both had large increases.—*PMB*

### Upcoming Communications Satellites

A new series of international and business communications satellites will be launched by 'workhorse' rocket systems, including updated Delta and Atlas/Centaur rockets, over the next few years. There is, of course, a 'long-shot' option that the space shuttle, as originally conceived, will be used to place the satellites in orbit, but no one is willing to bet right now that the shuttle will be functional and operational in

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time. Instead, the U.S. will employ updated versions of 15-20-year-old rockets to launch a series of satellites with names like 'INTELSAT,' 'INSAT,' 'Palapa,' and 'SBS' into geosynchronous orbits.

#### Communication Satellites

INTELSAT V is the first of a new generation of international telecommunications satellites sponsored by the 105-nation International Telecommunications Satellite Organization (INTELSAT), headquartered in Washington, D.C. The satellite, which weighs 1,928 kg at launch, has almost double the communications capability of early satellites in the INTELSAT series. It is positioned in geosynchronous orbit over the Atlantic Ocean so as to provide communications between North America and Europe.

INTELSAT V satellite is built by the Ford Aerospace and Communications Corp., Palo Alto, Calif., using system components developed by firms in France, the United Kingdom, the Federal Republic of Germany, Japan, and Italy. It has a capacity of 12,000 voice circuits plus two television channels.

The Atlas Centaur (AC-54) launch vehicle placed the INTELSAT V into a highly elliptical orbit from 166 by 35,788 km. It is from this orbit, at apogee, that a solid-propellant rocket motor attached to the satellite will be fired to circularize the orbit at geosynchronous altitudes over the equator. At that altitude, because the speed of the satellite in orbit matches the rotational speed of the earth, the satellite will remain in position over one spot.

This INTELSAT V launch costs approximately \$76.6 million, including \$34 million for the satellite and \$42 million for the Atlas Centaur and related launch services, which is reimbursable to NASA under the provisions of a launch services agreement signed in May 1980.

NASA Administrator, Robert A. Frosch, and Satish Dhanwan, secretary of the Government of India's Department of Space, signed an agreement recently in Bangalore, India, calling for the launch of two Indian communications/meteorological satellites. The two satellites, Indian National Satellite (INSAT) 1A and 1B, operating in geosynchronous orbit, will provide India with point-to-point voice and television communications, community broadcasting, and weather data. The two satellites are being built by Ford Aerospace Corporation.

The agreement calls for the Department of Space to be responsible for satellite manufacture and checkout and integration of the spinning solid upper stage (SSUS), which will boost the satellites into elliptical transfer orbit from which they will be maneuvered into their geosynchronous operating orbits. NASA will provide all other launch-related services through the Kennedy Space Center, Fla.

Stanley I. Weiss, NASA associate administrator for space transportation operations, and Dr. Suryadi, director general of posts and telecommunications for the Republic of Indonesia, signed an agreement in Jakarta, calling for the launch by NASA of two Indonesian communications satellites by January 1984.

The two satellites, Palapa B-1 and B-2, operating in geosynchronous orbit, will provide voice, video, telephone, and high-speed data services to Indonesia and other member states of the Association of Southeast Asian Nations—the Philippines, Thailand, Malaysia, and Singapore.

The agreement calls for Indonesia to be responsible for satellite checkout and integration of the spinning solid upper stage, and NASA will provide all other launch-related services through the Kennedy Space Center, Fla.

The first of three satellites in an advanced commercial communications system being established by Satellite Business Systems, McLean, Va., called SBS-A (SBS-1 in orbit), is being launched on a Delta launch vehicle. This is the 153rd launch of a Delta. Over the past two decades the McDonnell Douglas-built launch vehicle has had a mission success rate of better than 90%.

The launch will mark the debut of a new solid-fuel payload assist module (PAM-D), which will provide an approximate 20% increase in payload capability on missions to geosynchronous orbit over a Delta TC-364-4 third stage. Developed with private funding by the McDonnell Douglas Astronautics

Date	Mission	Launch Vehicle	Launch Site	Sponsor and Description
February	COMSTAR-D	Atlas Centaur	ESMC*	Comsat General Corp.—communications—reimbursable
March	INTELSAT V-B	Atlas Centaur	ESMC	INTELSAT—communications—reimbursable
March	GOES-E	Delta	ESMC	NOAA—weather—reimbursable
April	Navy 20 (Nova 1)	Scout	ESMC	DOD—transit—reimbursable
April	SBS-B	Delta	ESMC	Satellite Business Systems—communications—reimbursable
May	NOAA-C	Atlas-F	WSMC	NOAA—weather—reimbursable
June	INTELSAT V-C	Atlas Centaur	ESMC	INTELSAT—communications—reimbursable
June	RCA-D	Delta	ESMC	Radio Corporation of America—communications—reimbursable
June	FLTSATCOM-E	Atlas Centaur	ESMC	DOD—communications—reimbursable
July	Dynamic Explorer	Delta	ESMC	NASA—scientific
September	Navy 22 (Nova 2)	Scout	WSMC	DOD—transit—reimbursable
September	INTELSAT V-D	Atlas Centaur	ESMC	INTELSAT—communications—reimbursable
September	Solar Mesospheric Explorer	Delta	WSMC	NASA—scientific
October	RCA-C1	Delta	ESMC	Radio Corporation of America—communications—reimbursable
December	INTELSAT V-E	Atlas Centaur	ESMC	INTELSAT—communications—reimbursable

\*Eastern Space and Missile Center, Cape Canaveral, Fla.

†Western Space and Missile Center, Vandenberg Air Force Base, Calif.

Co., Huntington Beach, Calif., the payload assist module is the Delta version of the spinning solid upper stage designed for use in the space shuttle.

The SBS-A is a 550-kg satellite that will provide integrated, all-digital, interference-free transmission of telephone, computer, electronic mail, and video teleconferencing to SBS business and industrial clients. The service should be inaugurated early this year. The second satellite in the series is scheduled for launch on a Delta this year, and the third one will be launched from the space shuttle in late 1982. By 1983, SBS also plans to establish an intercity satellite telephone service that will connect up to 150 metropolitan calling areas.

SBS-A is a spin-stabilized satellite 216 cm in diameter, with a stowed height at launch of 282 cm. After deployment in its geosynchronous orbit at about 35,880 km above the earth, the telescoping solar panel cylinder will be extended and the communications antenna raised, giving the satellite an overall height of 860 cm. Each has a high-speed, all-digital 10-transponder system capable of relaying up to 480 million information bits of data per second, the equivalent of more than 10 million words. They are also the first U.S. domestic commercial communications satellites to use the higher, less congested 12/14 GHz (K-band) frequencies.

Once in orbit at 106°W over the equator—about due south of Santa Fe, N.M.—the satellite's antenna pattern will cover the continental United States, delivering higher power to metropolitan regions in the East, Midwest, and West Coast, where SBS customer communications traffic will be greatest. The payload assist module, being flown for the first time on Delta in place of the conventional third stage, is designed to inject the satellite into an elliptical transfer orbit ranging from a perigee, or low point, of 166 km to an apogee of 14,252 km (22,950 mi.). It is from this orbit, at the fourth apogee, that the SBS-A apogee kick motor is fired, which will place the satellite into its geosynchronous operating orbit.

#### The Launch Vehicles

Overall, Delta, in service since 1960, is 35.4 m tall and weighs about 192,098 kg at liftoff. The first stage is a long-tank derivative of the Thor vehicle, 22.5 m long and 2.4 m in diameter. Its main engine, burning RP-1 fuel and liquid oxygen, is rated at 920,777 N at sea level. It has a burn time of 3 min 43 s. First-stage thrust augmentation is provided by nine solid fuel strap-on motors that are 11.2 m long. Five of the motors are ignited at liftoff and four ignite after the first five burn out. Each motor, with a burn time of 57 s, provides an average of 379,298 N of thrust.

Delta's second stage, burning nitrogen tetroxide as the oxidizer and Aerozine-50 as the fuel, is 6.4 m long and 140

cm in diameter. It produces 43,592 N of thrust and burns for about 300 s. The second stage also contains the guidance system that generates steering commands for the first and second stages, as well as timing, staging, and engine re-starts when needed.

The Atlas Centaur is NASA's standard launch vehicle for intermediate weight payloads. It is used for the launch of Earth-orbital, Earth-synchronous, and interplanetary missions. Centaur was the nation's first high-energy, liquid-hydrogen/liquid-oxygen-propelled rocket. Developed and launched under the direction of NASA's Lewis Research Center, it became operational in 1966 with the launch of Surveyor 1, the first U.S. spacecraft to soft-land on the moon's surface.

Since that time, both the Atlas booster and Centaur second stage have undergone many improvements. At present, the vehicle combination can place 4536 kg in low Earth orbit, 1928 kg in a geosynchronous transfer orbit, and 907 kg on an interplanetary trajectory.

The Atlas Centaur, standing approximately 39.9 m high, consists of an Atlas SLV-3D booster and Centaur D-1AR second stage. The Atlas booster develops 1920 kN of thrust at liftoff, using two 822,920-N thrust booster engines, one 266,890-N thrust sustainer engine, and two vernier engines that develop 2890-N thrust each. The two RL-10 engines on Centaur produce a total of 133,450-N thrust. Both the Atlas and the Centaur are 3 m in diameter.

Until early 1974, Centaur was used exclusively in combination with the Atlas booster. It was subsequently used with a Titan II booster to launch heavier payloads into Earth orbit and interplanetary trajectories.

The Atlas and the Centaur vehicles have been updated over the years. Thrust of the Atlas engines has been increased about 222,400 N since their first use in the space program in the early 1960's.

The Centaur has an integrated electronic system that performs a major role in checking itself and other vehicle systems before launch and also maintains control of major events after liftoff. The system handles navigation and guidance tasks, controls, pressurization and venting, propellant management, telemetry formats and transmission, and inflates vehicle events.

The Atlas and Centaur stages of Atlas Centaur 54 arrived at Cape Canaveral Air Force Station August 6, 1980. The Atlas was erected on Pad B of Launch Complex 36 on August 12; the Centaur was erected on August 14. A terminal countdown demonstration test was conducted October 3 to verify the integrity of the vehicle-to-ground systems in an environment that duplicates launch conditions.—PMB

populations the size of most communities.

It is interesting to note that studies of the survivors of nuclear bombings of Hiroshima and Nagasaki showed that correlation between increases of cancer incidence and the degree of radiation exposure could be made only on a population basis. Individuals who had received significant radiation had recognizable chromosomal damage, but still, according to Koiata, those individuals with the greatest amount of damage were not necessarily those who got cancer. No increases in birth defects or miscarriages were observed statistically.

The assessment of chromosome damage is as much an art as a science. While blood cells must be carefully cultured, then stained and examined under the microscope. The 46 chromosomes in a human cell can be individually identified by their characteristic shapes and sizes. If there is damage, it often appears as breaks and deletions, or as rings, which are formed from chromosome fragments. Cells with damaged chromosomes usually die or repair the damage.

Although the chromosomes are the carriers of genes, almost never can specific chromosomal aberrations be associated with specific birth defects or cancer. One exception is Down's syndrome, in which individuals inherit an extra chromosome 21, and this extra chromosome shows up in all their cells. But most genetic defects and most DNA damage that may lead to cancer involve submicroscopic changes in DNA and quite often do not lead to physical changes in the chromosomes. There is only indirect evidence associating chromosome damage with birth defects and cancer. (Science, op cit).—PMB

#### Thermosphere Circulation Modeled

When solar storms force the earth's auroras to lower latitudes, winds in the thermosphere reverse direction and are whipped up to velocities of 2250 km/h. A computer model has now been developed that will describe the circulation of the thermosphere—a 400-km blanket enveloping the earth, with its bottom boundary at an altitude of 80 km—and its interaction with the auroras.

Raymond G. Roble, of the National Center for Atmospheric Research, explained at the AGU Fall Meeting that the thermospheric model might be useful to predict variations in storm time and atmospheric drag on some earth satellites. If the thermosphere's dynamics are better understood, he reasoned, more accurate predictions of a satellite's orbital decay can be made. The model may also help to predict the effects of communications equipment and magnetic forces on power grids.

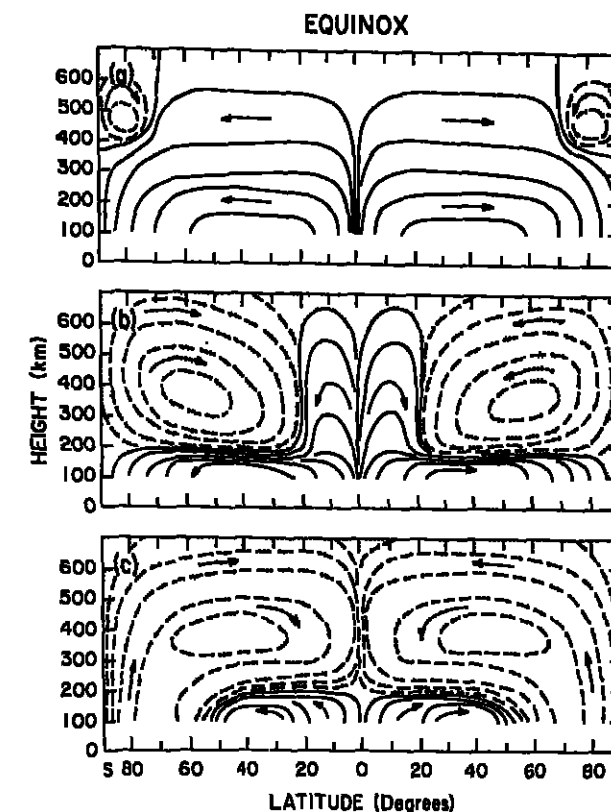
Developed by Roble, E. Cicely Ridley, and Robert E. Dickinson, the model is a set of meteorological equations adapted from the NCAR model of general circulation in the lower atmosphere. The model is constructed as a global grid of more than 60,000 points at 24 altitudes throughout the thermosphere. At each point, the equations calculate the dynamic relationships between temperature, pressure, winds, and other variables. Circulation patterns are computed by simulating progression of time.

The model incorporates a geomagnetic pole that tilts away from the geographic pole. Because auroras are centered around the geomagnetic poles, the tilt imparts a wobble to the daily circulation of the thermosphere in the auroral zones.

The thermosphere is heated continuously by ultraviolet radiation from the sun. The region's basic circulation moves from the hot daylight portions to the cool nightside and back, with winds blowing several hundred kilometers per hour. The mean circulation is from the equatorial region toward the poles.

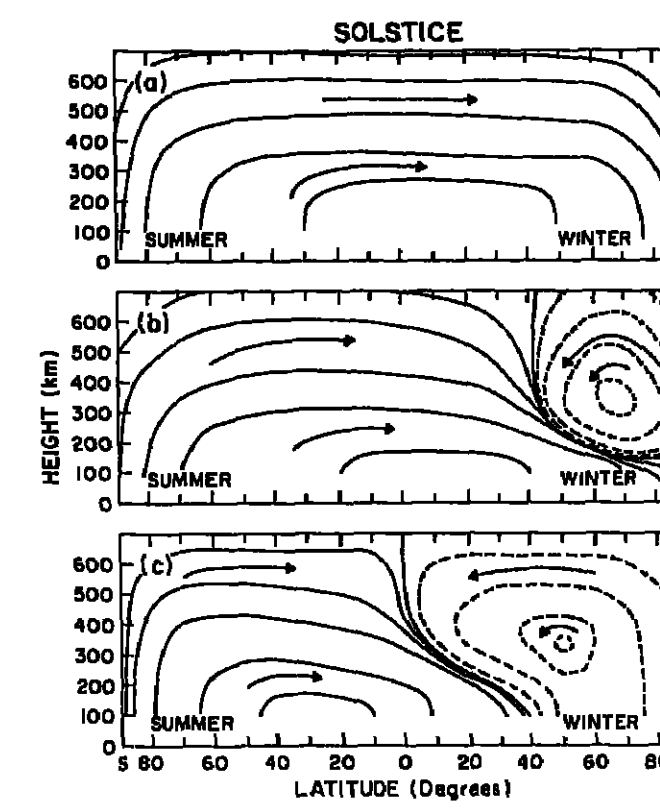
Sporadic auroral activity produces heat in the polar regions of the thermosphere, setting up an opposing circulation, for example, from the polar regions toward the equator. This

happens on a small to medium scale several times a day. During major geomagnetic storms, the aurora-induced circulation overwhelms its ultraviolet counterpart and reverses at



Schematic diagram of the zonal mean meridional circulation in the earth's thermosphere during equinox for various levels of auroral activity: (a) extremely quiet geomagnetic activity, (b) average activity, and (c) geomagnetic substorm. [Source: NCAR]

most the entire flow in the thermosphere. Winds in this powerful counterflow have been measured over 1600 km/h, and in one case over 2250 km/h. ☐



Schematic diagram of the zonal meridional circulation of the earth's thermosphere during solstice for various levels of auroral activity: (a) extremely quiet geomagnetic activity, (b) average activity, and (c) geomagnetic substorm. [Source: NCAR]

## Classified

EOS offers classified space for Positions Available, Positions Wanted, and Services and Supplies. There are no discounts or commissions on classified ads. Any type that is not publisher's choice is charged for at display rates. EOS is published weekly on Tuesday. Ads must be received in writing on Monday 1 week prior to the date of the issue required. Replies to ads with box numbers should be addressed to: Box \_\_\_\_\_, American Geophysical Union, 2000 Florida Avenue, N.W., Washington, D.C. 20008.

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**POSITIONS AVAILABLE**

**NSF.** The National Science Foundation, Division of Ocean Sciences, is seeking qualified applicants for the position of program director in the Physical Oceanography Program. The position is exempt from the competitive civil service. This appointment will be for 1-2 years. The program provides support for scientists primarily from academic institutions to pursue fundamental research in physical oceanography. The selected candidate will carry out program planning and budgeting, proposal evaluation, administration of research grants, and liaison with other federal agencies. Applicants should have a Ph.D. in physical oceanography, a related physical science, or the equivalent, plus at least 3 years of specialized experience in physical oceanography. Field experience in an academic institutional program is highly desirable. Salary range is from \$37,871 to \$50,112.50, depending on qualifications and experience. Those interested in being considered for the position should send letters of interest and CV to: Personnel Administration, 1800 G St., N.W., Rm. 212, Washington, DC 20560. Attn: E. Paul Bragg. For further information call (202)357-7841.

NSF is an equal opportunity employer

**Assistant Professor/Northern Illinois University.** Applications invited for a probable tenure track faculty position beginning August 1981. Candidates are being sought who are specialists in one or more of the following areas: Igneous petrology, economic geology, or mineralogy. The Ph.D. degree is required. The successful candidate will teach graduate and undergraduate courses and will be expected to pursue an active program of research in higher specialties.

Applications should include resume and the names and addresses of three persons who could serve as references. Inquiries and applications should be addressed to: L. D. McDevine, Search Committee, Department of Geology, Northern Illinois University, DeKalb, IL 60115. Application deadline: April 15, 1981. An equal opportunity/affirmative action employer.

**Research Faculty Position.** A faculty position at the research assistant professor level will be available at the Department of Geology, University of Miami from August 15, 1981 (this position will become a tenure-track position on August 1, 1982). Minimum qualifications are a Ph.D. in the geological sciences, a fair for teaching, and strong research interests as proved in publications. Areas of specialization (one or more of the following): geochemistry, economic mineralogy, petrology. Instrumentation available at the department: mass spectrometers for  $^{13}\text{C}/^{12}\text{C}$  and  $^{18}\text{O}/^{16}\text{O}$  analysis; mass spectrometers for rare gas analysis ( $^{40}\text{Ar}/^{39}\text{Ar}$  dating); a fully equipped radiocarbon laboratory with scintillation counting; atomic absorption units; thermoluminescence unit; video-taping and microcomputer systems; rock thin-sectioning laboratory; petrographic microscope; stereomicroscope, etc.

Preference will be given to candidates who would be able to contribute to the department's research program. A list of the equipment and computer facilities are available. The potential exists both for outside funding and for cooperative research. The successful candidate will be expected to conduct an active research program leading to publication. Applicants should submit a letter of application, resume, a copy of each transcript, and have three supporting letters sent to: Chairman, Department of Geology, University of Miami, P.O. Box 249172, Coral Gables, Florida 33124. Tel: (305) 284-4253. The University of Miami is a private, independent, international university. An equal opportunity employer.

**Meteorologist.** Sigma Data Computing Corp.'s Division of Information and Scientific Applications invites applications from meteorologists qualified to participate in an environmental modeling/chemical data assessment team effort. The applicant will evaluate atmospheric models and their data requirements to form a comprehensive multimedia modeling library system for assessment of toxic chemicals. The applicant will also provide recommendations for modification of existing model algorithms and R&D for anticipated continuing model development.

An M.S. degree or equivalent experience is a minimum requirement. Programming experience in FORTRAN and use or development of air quality models is desirable but not essential. Salary is commensurate with qualifications. Please submit resume and references to: Roger Long, Sigma Data Computing Corp., 2021 K Street, N.W., Suite 207, Washington, D.C. 20006.

**Sedimentologist.** The State University of New York at Binghamton has a vacancy for a sedimentologist at the assistant professor level. Candidates with research interests in sedimentology and a solid theoretical background are encouraged to apply. A Ph.D. with 0 to 5 years of teaching, research, and/or industrial experience is appropriate for the position. Salary is negotiable and competitive with academic institutions. The position is available in fall 1981. Please send resume and names of three references to: Chairman, Search Committee, Department of Geology, State University of New York at Binghamton, Binghamton, NY 13901. The State University of New York at Binghamton is an affirmative action/equal opportunity employer.

**Postdoctoral Research Associate/Mineralogy.** Applications are invited for research in high-resolution and analytical transmission electron microscopy of minerals and their analogues. Experience in crystallography, materials sciences, or electron microscopy is desirable. Send resume (including transcripts), statement of research interests, and names of three references to: P. R. Busck, Department of Geology, Arizona State University, Tempe, AZ 85281. Arizona State University is an EO/AAE employer.

**Sedimentary or Low Temperature Geochemist.** This is an assistant professor, tenure track position, although exceptional candidates of higher rank will be considered. We are looking for a geochemist to complement our strong programs in sedimentology, hydrogeology, organic geochemistry, and basin analysis. The teaching load is three courses per year—one beginning level geology course, an upper level geochemistry course, and a graduate course of higher choice. Introductory geology and summer field camp are also taught on a long-term rotating basis. A well-equipped laboratory and computer facilities are available. The potential exists both for outside funding and for cooperative research.

The successful candidate will be expected to conduct an active research program leading to publication. Applicants should submit a letter of application, resume, a copy of each transcript, and have three supporting letters sent to: Chairman, Department of Geology, University of Missouri, Columbia, Missouri 65211. The University of Missouri is an equal employment opportunity employer.

**Ocean Dynamicalist.** An academic position (tenure-track) for an ocean dynamicalist is presently available in the Department of Oceanography, Naval Postgraduate School (NAPPS/OSCOL). Present or ultimate research interest in areas of naval oceanography concern is desirable. Such areas include: ocean circulation modeling, especially prognostication on the oceanic synoptic scale; surface and internal gravity wave dynamics; synoptic analysis of oceanic data; and acoustical oceanography. The candidate should be willing and able to teach a variety of graduate courses in physical oceanography and related topics. The NAPPS/OSCOL has excellent computer, data archival, library, and research vessel facilities. The Department of Oceanography has close relations with the Fleet Numerical Oceanography Center, Naval Environmental Prediction Facility, and the Naval Laboratories. The department has a faculty of fifteen and a student body of 80 to 100. The overall emphasis is ocean prediction with present faculty and student research in coastal ocean, polar ocean, and air-sea interaction processes. The academic and research programs are conducted in close collaboration with the Departments of Meteorology and Physics. Salary will be determined by qualifications of the successful candidate. By January 1, 1981, please send a curriculum vitae, the statement of research and instructional interests to: Faculty Search Committee, Department of Oceanography, Naval Postgraduate School, Monterey, CA 93940. Visits by top candidates will be scheduled soon after. A decision will be announced by March 1, and the position should be occupied by about June 1, 1981.

The Naval Postgraduate School is an equal opportunity employer.

**Research on Lunar Samples.** Applicants for this postdoctoral research position should have experience in at least one of the areas: lunar-sample research, meteorite research, or neuroendocrinology. Salary about \$10,000 per annum. J. T. Watson, Institute of Geophysics & Planetary Science, University of California, Los Angeles, California 90024. UCLA is an affirmative action/equal opportunity employer.

#### Environmental Pollution, Chromosomes, and Health

In mid-May, 1980, President Carter declared a state of emergency at the Love Canal area, near Niagara Falls, New York. The reason for this was for the U.S. to underwrite the relocation costs (\$3-5 million) of some 2500 residents who, according to a report by the EPA (Environmental Protection Agency) may have suffered damaged chromosomes. These injuries were apparently caused by contact with toxic wastes that had been dumped in the area in the years prior to development for housing.

That the toxic compounds exist in the Love Canal and Niagara Falls subsurface zones, including public water supplies, appears to be established fact. That the residents of the Love Canal area suffered chromosomal damage may be established fact as well. Whether or not these two findings can be linked to ill health of the residents is another matter. Recently, the EPA report has been described as having "close to zero scientific significance," and has been "discredited" (Science, 208, 1236, 1980). The reasons for this disparity go beyond differences of opinion, beyond possible inadequacies of the EPA study, and even beyond problems that probably will arise from future studies, including those now in the planning stages. The problem is that even if victims have easily recognizable injuries from toxic substances (injury that apparently has not occurred to Love Canal residents), medical science usually cannot show a causal relationship. Even chromosomal damage is, at best, difficult to interpret. In ideal studies of significant populations and control groups, the association of toxic chemical to chromosome

damage and to cancer and birth defects is indirect and, up to now, has been shown to have little or no significance to an individual member of the exposed population.

Geophysicists concerned with groundwater resources and chemical pollution are becoming increasingly aware of the extent of such pollution caused by dumping of wastes. By the same token, residents of areas known to be polluted are becoming more concerned, and in some cases terrified. The residents of the Love Canal area have suffered, at least financially and psychologically, and the government has concluded that they deserve recompense. But, what of the real question of medical effects: cancer, miscarriages, birth defects, seizures, etc.? At this time, it would appear that the geoscientist concerned with pollution will have to proceed with studies, taking it on faith that uncontrolled disposal of toxic chemicals must cease.

The recent signing of the "superfund" legislation by President Carter will clear the way for release of \$1.6 billion for cleaning up sites that have been used as dumps of hazardous wastes. The residents of polluted or contaminated areas may find little solace for their injured emotional state. No doubt the long-term results of studies of the Love Canal dump site will be very beneficial, although perhaps not as direct as might be desired. In a short article on chromosome damage, G. B. Koiata (Science, 208, 1240, 1980) points out that while such damage can be an important result of exposure to toxic chemicals, some damage occurs naturally from numerous nontoxic causes. In fact, the normal number of cancer cases, birth defects (1% of all children born), and spontaneous abortions (as high as 50%) is so high that it is usually difficult or impossible to show significant increases, particularly in



**Drexel University/Atmospheric Scientist.** Three tenure track faculty positions are anticipated starting fall 1981. Applications are solicited from Ph.D.s with independent research experience in one of the following areas of atmospheric science: general circulation; climate dynamics with application in satellite meteorology; atmospheric optics, experimental or theoretical with emphasis in mesoscale probing; boundary layer turbulence modeling and atmospheric chemistry modeling. Rank and salary commensurate with experience. Send resume and references to Dr. William W. Eiden, Head, Department of Physics and Atmospheric Science, Drexel University, Philadelphia, PA 19104. An equal opportunity/affirmative action employer.

**Stable Isotope Geochemistry/University of Saskatchewan.** The Department of Geological Sciences has a vacant tenure track position at the assistant professor level for a stable isotope geochemist. Applicants should hold (or be about to receive) the Ph.D. degree, be qualified to instruct undergraduates in general geology and undergraduates and post-graduates in geochemistry and geology, be prepared to pursue a vigorous research program, and to establish contacts in the geochemistry research laboratory with Micromass 902 Double Collecting Mass Spectrometer. Letters of application, with curriculum vitae including the names of at least three referees, should be sent to W. G. E. Caldwell, Head, Department of Geological Sciences, University of Saskatchewan, Saskatoon, Canada, S7N 0W0.

**Sedimentary Petrology.** The Geology Department at the University of Vermont is seeking a sedimentary petrologist for a tenure track position at the assistant professor level. Research and teaching specialties should be in clastic sedimentary petrology with potential ancillary interests in petroleum geology, geomorphology, and hydrology. It is expected that the successful candidate will establish a field oriented research program which includes supervision of graduate (M.S.) and undergraduate students. A Ph.D. is required and teaching experience is highly desirable. The Geology Department at the University of Vermont is a seven member department having an M.S. program and a definite commitment to excellence in undergraduate education. Applications will be accepted until April 1, 1981. Candidates should send a resume and arrange for three letters of reference to be sent to John C. Drago, Academic Chairman, Department of Geology, University of Vermont, Burlington, Vermont 05405. The University of Vermont is an equal opportunity/affirmative action employer.

**Seismologist.** The University of Nevada Seismological Laboratory invites applications for the position of lecturer/research seismologist. Candidates with interest in seismological research related to earthquake hazard, earthquake prediction, theoretical source mechanisms, and/or seismic signal characteristics at regional distances are encouraged to apply. Teaching duties will consist of one undergraduate or graduate course per semester plus participation in graduate seminars. Ph.D. degree in geophysics with at least three years research experience in earthquake seismology is appropriate for this position. Salary up to \$30,000 for twelve-month contract, depending on background and experience. Position two thirds supported by state, one third by grants and contracts. Available on or after 1 May 1981. Deadline 1 March 1981. Candidates should send a letter of application, list of publications, statement of teaching and research interests, transcripts and names of three references to Alan Ryan, Director, Seismological Laboratory, University of Nevada, Reno NV 89597. AAEOE.

**Postdoctoral Research Associate.** Oceanographic Department of the Naval Postgraduate School seeks recent graduate to study the hydrodynamics through numerical ocean modeling of the physical-oceanographic processes active in the vicinity of the Arctic edge of Alaska. Problem areas include the effects of the complex bathymetry on the circulation and frontal formation, the dynamics associated with interlocking of water masses at the ice edge, and the mechanisms involved in ice retreat. Research will be performed in the context of an observational program which has acquired data and developed insights over the course of several years. Position starts March 1981 and is renewable annually. Salary depends upon qualifications. Send resume and the names and addresses of three references to Faculty Search Committee, Dept. of Oceanography, Naval Postgraduate School, Monterey, CA 93940. Equal opportunity/affirmative action employer.

**Staff Scientist/Ocean Margin Drilling Program.** Joint Oceanographic Institutions, Inc. (JOI), Inc. has immediate openings for two staff scientists to fill the positions of Field Programs Coordinator and Downhole Measurements Coordinator in its Ocean Margin Drilling (OMD) Science Program. Candidates having each of these positions report to the OMD Chief Scientist. They will be required to provide staff support to advisory committees in their area of concern, and will be responsible for implementing programs recommended by the OMD Science Advisory Committee, including oversight of the performance of individuals or groups under contract to JOI. Both positions require a Ph.D. in geophysics and a background in oceanographic research. The OMD Science Program is funded for FY 81. Initial appointment will be for a period of two years with the second year contingent upon the availability of funds. The positions may be filled on a non-exclusive basis. Salary will be commensurate. Send resume, statement of interest, and the names of three referees to: Thomas A. Dwyer, Chief Scientist, Ocean Margin Drilling Program, Joint Oceanographic Institutions, Inc., 2820 Virginia Ave., NW, Suite 610, Washington, DC 20037. The deadline for applications is February 28, 1981, or as soon thereafter as suitable candidates are found.

**Director/School of Meteorology.** The University of Oklahoma invites nominations and applications for the post of director of the School of Meteorology effective for the 1981-82 academic year. The school offers programs of study leading to B.S., M.S., and Ph.D. degrees in areas ranging from traditional atmospheric sciences to application-oriented climatology. Many of these programs have developed close, synergistic relationships with the activities of the National Severe Storms Laboratory, the Oklahoma Climatological Survey, and the OU-NOAA Cooperative Institute for Mesoscale Meteorological Studies. Relationships that offer attractive opportunities for innovative multidisciplinary and interdisciplinary programs.

Applicants should have a Ph.D. in meteorology or a closely related field and several years of relevant experience or equivalent qualifications, and should qualify for regular academic appointment. The University of Oklahoma offers a comprehensive meteorology program comprising about 120 undergraduates, 50 graduate students, 8 faculty members, and several research associates. The program has been highly productive as measured by its sponsored research activities and the success of its graduates. The director is expected to provide leadership that will sustain and improve the quality and character of meteorology at the University of Oklahoma as well as to contribute to the teaching and research programs of the school.

Nominations and applications should be sent to Wm. R. Upthegrove, Chairman, Meteorology Director Search Committee, 107 East Engineering Center, University of Oklahoma, Norman, Oklahoma, 73019.

Applications should include a resume, a list of publications, and names of at least three professional references. In addition, candidates are encouraged to submit supplemental statements of their professional goals and their impressions of the directions and goals for atmospheric sciences in 1980's. Initial screening will begin February 23, 1981; however, applications will be accepted and recruiting continued until the position is filled. The University of Oklahoma is an equal opportunity/affirmative action employer.

**Hydrogeologist.** The State University of New York at Binghamton invites applications for a permanent position in groundwater hydrology, starting fall 1981. It is desirable that applicant have teaching and research interests in one or more of the following: groundwater hydrology, modeling, flow through porous media, and environmental hydrogeology. However, applicants with interests in other areas will be considered.

Teaching responsibilities will include both undergraduate and graduate courses. The opportunity exists to initiate courses at all levels, but development of one lower-level undergraduate course is essential. Research facilities include: electron microscope, scanning electron microscope, X-ray diffractometer, atomic absorption and transmission spectrophotometers, and access to a large central computer as well as minicomputer in department. Appointment is planned as assistant professor, but not necessarily at beginning level. Salary is negotiable, but will be at competitive academic level. Applicants should submit resume and arrange for three letters of recommendation to be sent to James E. Sorrell, Chairman, Department of Geological Sciences, State University of New York at Binghamton, Binghamton, NY 13901.

State University of New York at Binghamton is an affirmative action/equal opportunity employer.

**Geochemistry/Brittle Deformation, University of New Brunswick.** The Department of Geology has a tenure track position available from July 1, 1981 at assistant professor or higher level. The successful applicant will be expected to teach both undergraduates and graduates as well as carrying out research and supervising graduate students. Applications will be accepted in the following fields: geochemistry of ore bodies, exploration, environmental or soil geochemistry, brittle deformation, rock mechanics or soil engineering. Applicants should have a Ph.D. and preferably postdoctoral experience. Applications including a curriculum vitae and names of three referees should be sent to P. F. Williams, Chairman, Department of Geology, University of New Brunswick, Fredericton, N.B. E3B 5A3.

**Structural Geologist/University of California, Santa Barbara.** Applications are invited for a tenure track appointment in structural geology to be filled during 1981-1982, subject to availability of funds. Rank dependent upon qualifications and experience, but preference will be given to the assistant professor level. Successful candidates must have a Ph.D. degree and strong desire and commitment to teach and direct M.A., Ph.D., and undergraduate students in both structural and field geology. He/she will be expected to develop a strong research program and obtain outside funding for its support. Additional duties may include teaching physical geology and supervising field geology.

Please send resume and evidence of teaching and research proficiency by March 31, 1981, and arrange for early submission of four letters of recommendation to Dr. Arthur G. Sylvester, Chairman, Department of Geological Sciences, University of California, Santa Barbara, CA 93106. (805) 961-3168. The University of California is an affirmative action/equal opportunity employer.

**Staff Scientist.** Staff Scientist to conduct research in paleo data analysis to understand environmental effects, in particular, to analyze oceanic paleo data to examine the composition of the atmosphere and its temperature variations, using 180, 360, 180, 600, 600, and CDC Cyber 175 computers. Require Ph.D. in physics or atmospheric science with solid knowledge of FORTRAN and JCL computer languages, and background in spectroscopy and computer simulation. Minimum one year background in research.

Send resumes to Rodney Smith, Manager of Staffing Systems and Applied Sciences Corporation, 8811 Kensington Avenue, Riverside, Maryland 20910. An equal opportunity employer.

**Meteorologist and Hydrologist/Saudi Arabia.** The School of Renewable Natural Resources, University of Arizona, invites applications for assignment as faculty members to the Institute of Meteorology and Arid Land Studies, King Abdulaziz University, Jeddah, Saudi Arabia. One year, renewable positions in meteorology and hydrology are available.

1. Ph.D. in meteorology with experience in undergraduate teaching and research. Curriculum includes courses in meteorological instruments and methods of observation, dynamic meteorology, synoptic meteorology, physical meteorology, and climatology.

2. M.S. in meteorology with practical experience in meteorological operations and undergraduate teaching. Knowledge of WMO procedures.

3. Ph.D. in a hydrologic science or engineering with experience in undergraduate teaching and in research. Major emphasis will be in the areas of surface and groundwater development, water management in an arid environment and in evaluating the hydrologic effects of development.

Description: The project is funded by the Saudi Arabian government through the U.S.-Saudi Arabian Joint Commission on Economic Cooperation. Administrative and logistic support is provided by the U.S. Treasury Department, while the program's implementation is by a contract with the Consortium for International Development. The goal of the project is to undertake technical cooperation to develop educational programs for meteorology, hydrology, and surface and environmental protection.

Salary and allowances: Highly competitive with 25% overseas adjustment, housing, car and other allowances.

Availability: February 1, 1981, or soon thereafter for spring semester; September 20, 1981, for fall semester. Initial appointment of one year or more contingent on performance.

Closing date: January 15, 1981 for spring semester; February 15 for fall semester.

Application: The application should include the following: (a) a letter detailing principal qualifications and interests; (b) a curriculum vitae; (c) name, address and telephone numbers of three references. Send to Martin M. Fogel, Director, CID/King Abdulaziz University Project, 317A Anthropology Building, University of Arizona, Tucson, AZ 85721, Telephone (602) 626-5544/3860. EEO/AAE employer.

**Program Manager/Meteorology.** Oceanographic Sciences, Inc., is seeking qualified applicants for the position of program manager for meteorological studies. Applicants should have an M.S. or Ph.D. in meteorology or atmospheric sciences, plus experience in the field. A broad general knowledge of air pollution, and an understanding of the air pollution regulatory environment, is helpful. Interested persons should send resume, references, and salary history to R. C. Banks, Oceanographic Sciences, Inc., 25 Cassilian Drive, Goleta, CA 93117.

**Remote Sensing/Ocean Engineering or Oceanography Faculty Position.** Applications are solicited for two permanent positions involving both research and graduate and undergraduate teaching. Ability to initiate funded research is desirable. Send resume, brief statement of research areas, and the names of three references to F. W. Morris, Search Committee, Department of Oceanography and Ocean Engineering, University of Virginia, Blacksburg, VA 24061. The University is an equal opportunity/affirmative action employer.

**Research Physicist.** Ph.D. and two years experience with ionospheric research related to communications properties or closely related area. Initial salary is \$21,000/year for 40-hour week. Interested applicants with these qualifications should call Mr. Bischoff at (301) 262-4400.

**Graduate Assistantships/Physics and Astronomy.** Graduate research assistantships and teaching assistantships in the Department of Physics and Astronomy of the University of Iowa are available to well-qualified students. The department has vigorous research programs in space physics, plasma physics, acoustics, astronomy, astrophysics, physical physics, elementary particle physics, laser physics, nuclear physics, and solid state physics. Assistantships can begin in June, August, or January. Please address your inquiry to Department of Physics and Astronomy, The University of Iowa, Iowa City, IA 52242.

**Institute of Space and Atmospheric Studies/University of Saskatchewan.** Applications are invited for postdoctoral research positions in solar physics and atmospheric dynamics. Term or full-time appointments are available. Experimental ability or experience with optical or radio techniques is desirable. Work may involve rocket, balloon or observational measurements and their interpretation. Send resume, references and research interests to: D. J. McEwen, Institute of Space and Atmospheric Studies, University of Saskatchewan, Saskatoon, Canada S7N 0W0.

**Faculty Position.** The Department of Geology of the University of New Mexico seeks applicants for a position in clay mineralogy, low-temperature geochemistry, carbonate petrology, or economic geology. The appointment may be at the assistant, associate or full professor level contingent on approval of funding from the university. The individual must have a Ph.D. degree and strong desire and commitment to teach and direct M.A., Ph.D., and undergraduate students in both structural and field geology. He/she will be expected to develop a strong research program and obtain outside funding for its support. Additional duties may include teaching physical geology and supervising field geology. Please send resume and evidence of teaching and research proficiency by March 31, 1981, and arrange for early submission of four letters of recommendation to Dr. Arthur G. Sylvester, Chairman, Department of Geological Sciences, University of California, Santa Barbara, CA 93106. (805) 961-3168. The University of California is an affirmative action/equal opportunity employer.

**Associate Director/Marine Science Institute.** The University of Texas at Austin seeks to fill the open position of associate director of the Marine Science Institute. The associate director is responsible for research and intellectual leadership of the Institute's Galveston Geophysics Laboratory. The position carries the line responsibility of senior administration for the Galveston Geophysics Laboratory. Duties include research planning, management, fiscal monitoring and budgeting, personnel review and assignment, coordination of scientific programs and ship operations, administrative supervision, liaison with industrial and agency sponsors, representation and other directorship duties.

The Galveston Geophysics Laboratory maintains modern computing facilities, research laboratories, and two deep-ocean research vessels, the R/V Fred Moore and the R/V Ida Green. Research at Galveston includes programs in marine geophysics, marine geology, solid earth geophysics, earthquake and extra-terrestrial seismology, and instrument systems design, both basic and applied.

Applicants are asked to send the following: (1) Vita—including list of publications. (2) Brief statement on current research and support. (3) Brief statement on administrative experience. (4) Brief statement on teaching experience. (5) Names of six persons who may be contacted for personal and professional recommendations.

A letter of application and the above requested information should be sent to:

Dr. Robert Moore, Director  
Marine Science Institute  
University of Texas  
P.O. Box 7989, University Station  
Austin, Texas 78712

Salary based on qualifications. Ph.D. required. The successful candidate will also be considered for tenure appointment in the Department of Marine Studies. Position to be filled as soon as possible. Early application advised. Position located in Galveston, Texas.

An equal opportunity/affirmative action employer.

**Faculty Positions/Astronomy/Space Physics.** The Department of Astronomy of Boston University invites applications for one or two tenure track faculty positions opening September 1981. Emphasis will be placed on active research experience as well as interest in graduate and undergraduate teaching. We are considering good candidates from any field of astronomy or space physics. Applicants should send resumes and the names of three references to M. D. Paggiannini, Department of Astronomy, Boston University, Boston, MA 02215. Boston University is an equal opportunity employer.

**Virginia Polytechnic Institute and State University.** Ignite Petrology and Geochemistry/Research Associate. Origin and tectonic significance of granitic rocks. Project involves petrography, analytical chemistry, mineral chemistry, isotopic studies, and field mapping. Send resumes to: Dr. R. W. Woodsworth, Chairman, Department of Geological Sciences, Virginia Poly. Inst. and St. Univ., Blacksburg, VA 24061. The University is an equal opportunity/affirmative action employer.

**Structural Geologist.** The Department of Geosciences of Purdue University invites application for a tenure track faculty position in structural geology, starting in August 1981. Rank and salary will be commensurate with qualifications. A Ph.D. is required. The individual will be expected to teach undergraduate and graduate courses in structural geology and tectonics, participate in summer field courses, and pursue an active research program. Preference will be given to a candidate with an applied field orientation and a strong background in the quantitative analysis of field data. The department has active programs in petrology, geophysics, and engineering geology and has a close working relationship with the geological group in civil engineering and the Laboratory for Applications of Petroleum Geology. Closing date for applications is April 1, 1981. Applicants should send a resume, the names, addresses, and telephone numbers of three referees, and a brief statement of research interests to R. H. McCullister, Department of Geosciences, Purdue University, West Lafayette, IN 47907. Purdue University is an equal opportunity/affirmative action employer.

**Synoptic/Dynamic Meteorology.** Description: The Geophysical Institute and Division of Geosciences, University of Alaska, invite applications from qualified scientists for a full-time (12 month) faculty position at the Assistant or Associate Professor level. The successful candidate will be expected to prepare and submit research proposals to external agencies and to work cooperatively with ongoing research programs. He/she will be also expected to teach occasional courses in synoptic/dynamic meteorology at the upper division and graduate levels.

Qualifications: Ph.D. in Meteorology. Research experience in advanced analysis and diagnostic studies of global-scale meteorological processes is essential, preferably over the full height of the atmosphere (0-100 km). Preference will be given to applicants who can utilize their expertise in synoptic/dynamic meteorology to synthesize the results of various ongoing research projects in mesoscale and large-scale meteorology, cloud physics, radiation, aeronomy, and space physics into a better understanding of the large-scale meteorology of the North Pacific and polar regions. Teaching experience at the undergraduate and graduate levels is desirable. Salary: \$43,500 (Assistant Prof.) to \$54,000 (Associate Prof.) per year, dependent upon qualifications and experience. Application: For further information, including recent annual research report, write to Director, Geophysical Institute, University of Alaska, Fairbanks, AK 99701. Closing date for applications is February 28, 1981. The University of Alaska is an equal opportunity/affirmative action employer.

## Meetings

### Geophysical Fluid Dynamics

The symposium on geophysical fluid dynamics, part of the European Geophysical Society's 8th meeting in Uppsala, Sweden, August 24-29, will include special sessions on the physics of lakes and fjords.

The special sessions will include discussion of circulation and the effects of the earth's rotation; seasonal and climatic effects; stratification; heat, momentum, and gas transfers; surface and internal waves; tides; effects of islands, bays, and sills; effects of river or melt water inflow; ice; heat flow; and sedimentation.

Potential contributors should notify the convenors by April 30. The convenors are J.-E. Weber, Institute of Geophysics, University of Oslo, P.O. Box 1022, Blindern, Oslo 3, Norway, and S. A. Thorpe, Institute of Oceanographic Sciences, Brook Road, Wormley, Godalming, Surrey, England GU8 5UB.

Deadline for receipt of abstracts is June 1. Abstracts should be sent to K. M. Storevold, Program Committee Chairman, Universitetet i Bergen, Geofysisk Institutt, Adv. C, Allegt, 70, N-5014 Bergen-Universitetet, Norway. \$

### Environmental Systems Conference

A call for papers has been issued for a working conference entitled Environmental System Analysis and Management. Sponsored by the International Federation for Information Processing (IFIP), the conference is scheduled for September 28-30 at the IBM Scientific Center in Rome.

## GAP

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### Aeronomy

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6111 SCATTERING OF RADIATION BY PARTICLES AND WAVES  
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6120 SCATTERING OF RADIATION BY PARTICLES AND WAVES

### Geochemistry

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1411 CHEMISTRY OF THE ATMOSPHERE  
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The conference is intended to provide an international forum for a broad interdisciplinary exchange of views among scientists who work in environmental science. Main areas of interest include water resources planning and management; air, water, and soil pollution; natural resources management; urban and regional planning; food and agriculture; energy and environment; environmental data bases; information systems for environmental problems; and environmental systems telemonitoring.

Abstracts of about 1000 words should be submitted in triplicate by March 15. Full papers will be required by September 30. Address abstracts and other conference correspondence to S. Rinaldi, Centro Teoria del Sistem, CNR, Politecnico, Via Pontio, 34/5, 20133 Milano, Italy; telex 333467 or telephone 02-2367241. \$

### Mexican Geophysical Meeting

Abstract deadline for the 1981 meeting of the Union Geofisica Mexicana is March 30, 1981. The meeting is scheduled for May 6-9 in Manzanillo, State of Colima, on the Pacific coast of Mexico.

Sessions will be held on the physics and chemistry of the earth's interior, exploration geophysics, atmospheric sciences, physical oceanography, and space and planetary physics.

Registration for the meeting is \$33 for UGM members, \$11 for students, and \$65 for nonmembers. Nonmembers wishing to join UGM may apply for membership and should include the annual fee (\$30 for active membership, \$30 for associate membership, or \$9 for student membership).

For more information about the meeting or about joining the Union, write to Union Geofisica Mexicana, Comité Reunión 1981, Instituto de Geofísica, UNAM Cd. Universitaria, Mexico 20, D.F. \$

## Senior Position in Earth Science

The Earth Sciences Division of the LAWRENCE BERKELEY LABORATORY has several comprehensive research programs involving the earth sciences. An opening exists for a person with an established national reputation in a scientific discipline in Earth Sciences, preferably geomorphology or hydrogeology, to assume a position of responsibility for the University of California, and director of major research programs such as concerned with radioactive waste storage.

Duties will include taking the scientific initiative and direction and management of ongoing projects, including the nuclear waste isolation field involving more than 30 scientists and engineers at LBL and collaborative work with several academic and research organizations. Additionally, the position involves establishment of emerging programs, expansion of research facilities and pursuit of new areas of investigation.

The successful candidate should have extensive experience and proven capabilities in directing and achieving programmatic goals of complex research projects involving teams of senior scientists and engineers. A Ph.D. in a field of the Earth Sciences is preferred with significant applicable experience. Salary: over \$50k.

Applications will be considered no later than April 1, 1981. Interested individuals should forward two resumes including salary history to: Employment Office, LAWRENCE BERKELEY LABORATORY, One Cyclotron Drive, Berkeley, CA 94720. An equal opportunity employer M/F.

**LAWRENCE BERKELEY LABORATORY**

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(News cont. from page 33)

close encounter with Saturn and Titan in November 1980. Voyager 1 has achieved the prescribed Saturn/Titan scientific objectives. The assessment of the health of the Voyager 2 spacecraft and instruments indicates that there is a reasonable probability that the 5-year journey to Uranus can be on schedule and a scientifically productive flyby can be achieved. "On this basis, therefore, the decision has been made to retain the present Uranus trajectory for Voyager 2," said Andrew J. Stofin, acting associate administrator for space science. He pointed out that retargeting of the spacecraft to provide another close Titan flyby could have been made as late as early 1981.—PMB

### Mutch Memorial Plaque Unveiled

A plaque commemorating Thomas A. Mutch, former associate administrator of NASA's Office of Space Science, was unveiled in a ceremony at the National Air and Space Museum earlier this month. Mutch was lost while mountain climbing in the Himalayas in October (EOS, October 28, p. 693). The plaque will be affixed to the Viking 1 lander, renamed the Mutch Memorial Station, during a future Mars mission.

A follow-up Mars mission has been suggested for the 1990's, and although no funding is available now, there is talk, NASA says, of sending a roving spacecraft to Mars that would allow the plaque, scoop up some Martian terrain, and bring the sample back to earth. Until the plaque can be transported to Mars, it will remain at NASA headquarters in Washington, D.C.

At the same ceremony, early in January, NASA accepted a \$60,000 check from The Viking Fund, a private organization under the auspices of the American Astronomical Society. The check represents individual contributions to support the continued operation and scientific analysis of the Viking 1 lander on Mars.

To seek additional support for Viking, NASA has designated July and August, the fifth anniversaries of the Viking arrival at Mars, as Viking Fund months. The Fund's donation will pay for the acquisition by the NASA Deep Space Network of data transmitted by the Viking lander during those months. □

### Science Education Research Program

A deadline for receipt of research proposals on science literacy and science, technology, and society has been set by the National Science Foundation's Research in Science and Education (RISE) program. March 9 is the target date set by NSF to insure that proposals are considered for the RISE fiscal 1981 budget, which is expected to total \$6 million.

RISE's purpose is to examine the science literacy of the U.S. public and to determine the public's needs. Although schools have been responsible for teaching science, only 50% of the American public receive formal science instruction after 15 years of age, according to NSF. Those who do not receive formal training must rely on a combination of electronic and print media, museums, and public agencies for science information.

For additional information and RISE guidelines on the preparation of formal proposals, contact the Program Director, RISE/SEDR, NSF, 1800 G Street, N.W., Washington, D.C. 20550, or call 202-282-7745. □

### New Map Data Catalog

Map byproducts, including aerial photographs, color separations, map data in computer form, and other materials used in or produced during mapping, are described in a new catalog published by the U.S. Geological Survey.

The 48-page hardcover catalog is the first listing of the unpublished USGS civilian cartographic holdings. It covers such items as mapping photographs, computer-enhanced LANDSAT pictures of Earth, cartographic data in computer form, microfilm and microfiche records, and a variety of features, including color separations, made in compiling and printing maps. The catalog also describes out-of-print maps available from USGS, along with land-use and land-cover maps, and other unusual items, such as slope maps and orthophotocopies. The catalog explains how to order advance copies of maps before they are published.

"Map Data Catalog" is available for \$3.50 from the USGS, Branch of Distribution, 604 Pickett Street, Alexandria, Virginia 22304. Orders must include check or money order, made payable to USGS. A color poster that summarizes the contents of a catalog, "MiniCatalog of Map Data," is available free upon request from NCIC, USGS, 507 National Center, Reston, Virginia 22092. □

### Geophysicists

J. C. I. Dooge, secretary of the Royal Irish Academy, has been elected secretary general of the International Council of Scientific Unions. A member of the International Union of Geodesy and Geophysics, he is the head of the civil engineering department at University College in Dublin.

M. F. Merer, president of the International Association of Hydrological Sciences, has been made an honorary member of the International Glaciological Society.

Thomas E. Pyle has been appointed deputy director of the National Ocean Survey. He was formerly head of the Office of Naval Research's Marine Geology and Geophysics Program and director of the ONR Detachment Washington Liaison Office.

Harold C. Urey, 87, a major contributor to the development of the atomic bomb, died January 6 in La Jolla, California. In 1934 he was awarded the Nobel Prize for chemistry for his discovery of deuterium. Urey was the director of the atomic bomb project at Columbia University during World War II. He had been professor-at-large at the University of California since 1958. Urey was an honorary fellow of AGU.

## New Publications

### Quantitative Seismology, 1, Theory and Methods

K. Aki and P. G. Richards, W. H. Freeman, San Francisco, xiv + 557 pp., 1980, \$35.00.

Reviewed by Freeman Gilbert

*Quantitative Seismology*, by Aki and Richards, will find a prominent place in the library of every seismologist. The two-volume work, of which the first is reviewed here, treats seismology as a branch of physics with a well-defined theoretical basis coupled with an observational program providing data of high quality. Very roughly, the first volume is devoted to the theoretical basis of seismology and the second to data analysis, interpretation, and problems of inference.

After a brief introductory chapter, the authors devote chapter 2 to the basic elements of the theory of elasticity. The conservation equations for linear and angular momentum are derived, and the classical constitutive relations are introduced. The concept of superposition for linear systems is introduced, and the Green's functions notation is used to derive compact representation theorems.

Dislocation sources and volume sources are introduced in chapter 3, and radiation from a point source is discussed in chapter 4. Here we meet P waves and S waves for the first time. A generalization of the far field expressions for homogeneous media leads to a discussion of ray theory in heterogeneous media, which is followed by a discussion of radiation patterns of body waves in a radially stratified medium.

The authors use the technique of introducing their mathematical methods in small doses while always emphasizing the physical meaning of their results. As a consequence, each succeeding chapter is mathematically only slightly more difficult. The procedure is an effective one and permits the authors to adopt an economical style without sacrificing either continuity or content.

Chapters 2-4, with their basic theorems and concepts, make the transition to boundary value problems a smooth one. The reflection, transmission, and conversion of plane P and S waves at a plane discontinuity are treated in chapter 5. Inhomogeneous plane waves are introduced, and the basic properties of Rayleigh and Stoneley interface waves are derived. Chapter 5 closes with a brief discussion of the effects of attenuation and anisotropy.

Chapter 6 is the last, most difficult, and most interesting of the introductory chapters. It is devoted to Lamb's problem, the problem of the interaction of cylindrical and spherical waves with a plane interface. Here, the aspiring theoretical seismologist cuts his teeth. The classical approach of the Weyl and Sommerfeld integrals is developed and approximate results derived via steepest descents. The exact solutions, obtained by the operational methods of Cagniard, de Hoop, and Pekeris, are then presented. The reader is exposed to a detailed study of the problem and its methods of solution. Complex variable theory and contour integration are used extensively but always with an eye to the physical interpretation of the results. Consequently, the reader is presented with new insight and understanding of diffracted pulses, head waves, interface pulses, and leaking waves.

It is the authors' intent that chapters 2-6 be introductory in nature, a sort of prologue to the heart of the first volume, chapters 7-9. They have been successful. It is quite evident that considerable care and effort have gone into the structure and content of chapters 2-6. Having assimilated the material therein, the reader is prepared for the following chapters on surface waves, free oscillations, and body waves.

The propagation and dispersion of surface waves is the topic of chapter 7. The concepts of phase velocity and group velocity are introduced by the use of the method of stationary phase, and the relation between spatial and temporal attenuation is derived. The bulk of the chapter is devoted to the basic boundary value problem for a stratified half-space.

Both the ODE approach and the variational approach are described. In the ODE approach, the authors expound the popular methods of numerical integration, the Thomson-Haskell matrix method, and the method of minors. The variational method is used to derive functional derivatives of phase velocity with respect to elasticity and density and to elucidate the Rayleigh-Ritz method for computing eigenvalues and eigenfunctions. The chapter concludes with Rosenbaum's classic theory of leaky modes. Given the identity and reputation of the authors, it is no surprise that this chapter on surface waves is up to date and very well written. It provides the reader with the knowledge and the methods to approach a research problem in this important branch of seismology.

Chapter 8, on free oscillations, could have preceded chapter 7 with some advantage to the logical structure of the text. In this way the transition from free oscillations to traveling waves in a spherically stratified medium to surface waves in a plane stratified half-space would appear in an orderly manner. It is a small point and detracts not at all from the quality of the book.

After deriving the Lagrange-Rayleigh excitation formula for the normal modes of a mechanical system, the authors introduce vector spherical harmonics and show the basic decomposition for a stratified sphere into spheroidal and toroidal modes. The effect of self-gravitation is included in the derivation of the governing ODE for free oscillations. The eigenvalue problems here are very similar to the ones in chapter 7, and they are solved with similar techniques, the two most used ones being  $n$ th order, one step methods for the ODE and the Rayleigh-Ritz method. Some observational results, principally for the Colombian earthquake of July 31, 1970, are presented to illustrate the methods used in very long period seismology. The chapter closes with a brief discussion of splitting caused by the rotation of the earth.

Chapter 8 is a very good introduction to the subject. It is basic material that must be mastered by anyone desiring to become a research worker in low frequency seismology. The growth in the subject has been very rapid in the past decade, so much so that a separate text could be devoted to it.

The propagation of body waves is the subject of chapter 9. It is easily the most technically demanding chapter in volume 1. The heterogeneity of the earth, particularly its major discontinuities in structure, leads to some challenging problems in the branch of seismology embraced by body waves. Classical ray theory, first discussed in chapter 4, is extended in chapter 9, and the reduced travel time, the integral over depth of the vertical slowness, is introduced. This variable, commonly named the tau variable, is ubiquitous in seismology generally and plays a central role in the present chapter.

Both plane layered media and smoothly stratified media are considered in detail. For the former, the operational method of Cagniard and others as well as the reflectivity method are presented. For the latter, WKBJ theory and the partial wave expansion are utilized. There are numerous examples to illustrate the methods. For instance, the frequency dependence of diffracted P waves is discussed by way of the Watson transformation (really due to Cauchy), and generalized PKP waves form the subject of the whispering gallery effect. In every case, the presentation is well motivated and is clearly stated.

Volume 1 closes with a chapter on seismometry. The standard types of seismographs are discussed and their response equations derived. Seismic accelerations range from 1 g in the epicentral area of some earthquakes to  $10^{-11}$  g or less for free oscillations excited by a moderate ( $M_s = 6.5$ ) earthquake. Several types of seismographs are needed to cover such a very large range of signal amplitudes, and chapter 10 describes them in enough detail for the reader to grasp the basic ideas. Modern seismometry is a large subject and could support a textbook quite easily.

*Quantitative Seismology* is a very successful book. It is well designed for teaching a graduate course in theoretical

seismology and is destined to become the standard reference on the subject. There is an extensive bibliography, a well prepared index, and a variety of figures, each carefully prepared, well captioned, and coordinated with the text. Each chapter is followed by a well chosen set of illustrative problems. Aki and Richards have done a great service for the rest of us.

Freeman Gilbert is with the Institute of Geophysics and Planetary Physics, Scripps Institution of Oceanography, La Jolla, California.

### Numerical Modeling of Marine Hydrodynamics

H.-G. Ramming and Z. Kowalk, Elsevier Oceanogr. Ser., Elsevier, New York, xii + 368 pp., 1980, \$63.00.

Reviewed by R. W. Garwood, Jr.

At the invitation of the Polish Academy of Science, Hans-Gerhard Ramming presented a series of lectures entitled "Numerical methods and their applications in shallow water areas" during April 1977 in Gdansk. Zygmunt Kowalk has assisted Ramming in assembling this lecture material and in combining it with some supporting theoretical subject matter into book form and in translating it into English.

To date there exists no textbook on numerical modeling of dynamical processes in the ocean, and any contribution that could fill this gap would be welcome. The rather promising title is misleading because it implies a more general treatment of ocean dynamics than is in fact presented. A more fitting title would have properly emphasized the primarily shallow-water applicability of the subject matter.

The list of numerical methods broached is not exhaustive, but the reader is exposed at least briefly to many of the standard techniques: finite differencing, stability and convergence criteria, iterative methods, physical versus numerical solutions, nonlinear methods, filtering, explicit and implicit schemes, and Galerkin methods. Throughout the text, numerical techniques are introduced informally and only as required in solving particular dynamical problems. The authors frequently invoke the adjective "well known" in referring to numerical techniques as they are introduced in the text. Even though the development is frequently cursory, an extensive list of up-to-date references is provided at the end of each chapter.

The chapter and section format is based upon dynamical topics rather than upon numerical techniques, but no strong physical foundation is laid. Although the first chapter does present the basic equations of motion, it appears to have been added as an afterthought. Lack of any mention of thermodynamics at this juncture tips the reader off to the limited

utility of the contents of the following chapters: Stratification and buoyancy effects will not be treated in depth. There are brief discussions of the effects of known (measured) density structure upon baroclinicity and upon the vertical exchange of momentum, but no consideration is given to modeling of the temperature and salinity fields.

The second and third chapters deal with steady motion: first some general numerical techniques for solving steady state systems of equations followed by a collection of some specific steady state problems in hydrodynamics. These problems range from classical ones such as Stommel's wind-driven circulation to a presentation of more recent developments in the understanding of turbulent boundary layer flows involving second-order closure using the turbulent kinetic energy budget.

Unsteady problems and accompanying numerical methods are combined into a single chapter. Again, little attention is given to stratification. The lack of a section on mixed layer modeling is a notable omission even for a treatise on shallow-water dynamics.

The next two chapters on tidal models in ocean basins, coastal zones, estuaries and rivers, and the following chapter

on the modeling of diffusion and dispersion of pollutants are the high points of this book. Clearly, it is the pursuit of the applied aspects of these general topics that is most interesting to the authors and gives rise to the expanded treatment here.

Even though the text is typewritten and is occasionally disjointed because of the English translation, it is quite readable, and typographical errors are few.

In conclusion, this text is best suited to those who are already versed in dynamical oceanography and who have some experience with numerical methods as well. It does help to fill the gap in readily available material on numerical modeling in oceanography, and it can be a useful addition to the reference library of any modeler of ocean dynamics. However, the value of this book as a basic textbook is not comparable to analogous texts in meteorology. Perhaps this is being too critical of a volume that was never intended to be more than a compendium of two scientists' salient experiences in hydrodynamical modeling.

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An M.S. degree or equivalent experience is a minimum requirement. Programming experience in FORTRAN and use of development of water quality models is desirable but not essential.

## NRL NAVAL RESEARCH LABORATORY has several vacancies for EXPERIMENTAL PHYSICISTS and MATHEMATICIANS (In Underwater Acoustics) GS-9 through GS-12 \$18,500 — \$35,000 per annum (Commensurate with Qualifications & Experience)

**Positions Involve:**  
Planning and conduct of experimental and theoretical Ocean Acoustic Investigations. Positions will require experimental design and conduct leading to the acquisition of surface scattering and bottom reverberation data. Involves direction of the reduction of data into temporal, spatial, and spectral components, comparison of results with existing theories, model validations, and extrapolation of results to other problem areas. Knowledge of applied mathematics, Fourier and spectral analysis techniques, and FORTRAN programming is desirable.  
(Full time positions #A-5100-15-EF) OR

**Positions Involve:**  
Investigation of the temporal and spatial variability of the acoustic signal field, beamformed array signal data, and signal spreading in angle, using a mini-computer signal processing system. Knowledge of spectral analysis, data smoothing, Fourier transform theory, wave propagation, array/antenna performance, and FORTRAN programming is desirable.  
(Full time positions #B-5100-15-DD)

trips at sea. An advanced degree or equivalent experience beyond the bachelor's level in physics and/or mathematics is required.

These positions are in the Federal Career Service with all Civil Service benefits. Travel and transportation expenses may be paid for selected applicants. Interested applicants should send a detailed resume or Personal Qualifications Statement (SF-171), with appropriate position number to:

Naval Research Laboratory  
Civilian Personnel Office  
Code: 5160-15 EOS  
4555 Overlook Avenue, S.W.  
Washington, D.C. 20375  
An Equal Opportunity Employer U.S. Citizenship Required

**Graduate Assistantships/Physics and Astronomy.** Graduate research assistantships and teaching assistantships in the Department of Physics and Astronomy of the University of Iowa are available to well-qualified students. The department has vigorous research programs in space physics, plasma physics, acoustics, astronomy, astrophysics, atomic physics, elementary particle physics, laser physics, nuclear physics, and solid state physics. Assistantships can begin in June, August, or January. Please address your inquiry to Department of Physics and Astronomy, The University of Iowa, Iowa City, IA 52242.

**Space Plasma Physicist.** The space plasma theory group at UCLA has begun a program of numerical simulations of problems in space plasma physics. We seek a Ph.D. level theoretical plasma physicist with a record of accomplishment in plasma numerical simulations, as evidenced by publications, who is interested in the theoretical research in magnetospheric physics. The applicant should have extensive experience in the design and implementation of plasma simulation codes, either kinetic or fluid. Prior experience with simulation of space plasmas is highly desirable. The position will be available September 1, 1981. Salary commensurate with experience and qualifications. Interested parties should send a curriculum vitae, bibliography and names of three suggested referees to Professor P. J. Coleman, Institute of Geophysics and Planetary Physics, UCLA, Los Angeles, California 90024, with a copy to K. Chwin at the same address.

UCLA is an affirmative action/equal opportunity employer.

**Personalized Estwing Pick Hammer.** 22 ounce, fully polished head with your name branded in leather handle and sheath. Excellent award gift or personalized item for field work. Special offer on orders of two or more hammers for \$18.00 each (\$20.00 ordered singly). Specify up to 20 letters & spaces. Call or write now. Western Heritage, 101 S. Washington St., Hinsdale IL 60521, tel (312) 864-5228.

**Services.** The complete Geophysical Year last appeared in the December 30 EOS.

**Boldface indicates meetings sponsored or cosponsored by AGU.**

**1981**

Mar. 9-12 ERRSAC/NOAA Land Remote Sensing Applications Conference, Danvers, Mass. Sponsors, NASA,

Apr. 13-15 AGU Spring Meeting in Baltimore, are to determine the availability of data needed for correlative IMS studies; to examine the scientific results of the various workshops; to analyze the strengths and weaknesses of the various workshop forms; and to identify the work required to fulfill overall IMS objectives.

The first day of the symposium will examine the data available for correlative purposes. Of greatest interest are data sets that cover a large fraction of the IMS time period (1977 to 1979). The second day will concentrate on the examination of scientific results from previous IMS workshops. On the final day, the scientific progress of the IMS analysis will be examined.

Further details about the symposium can be obtained from the convenors: C. T. Russell, Institute of Geophysics, University of California, Los Angeles, California 90024, and D. J. Southwood, Physics Department, Imperial College, London SW7, United Kingdom. \$8

**Four copies of paper summaries are due February 13.** Summaries should be between 500 and 1000 words. Authors will be notified of paper selection by March 18. A camera-ready copy of the manuscript will be due June 26.

Send summaries to Douglas B. Morrison, Purdue University/LARS, 1220 Potter Drive, West Lafayette, Indiana 47906. \$8

**Processing Remotely Sensed Data**

A call for papers has been issued for the Seventh International Symposium on the Machine Processing of Remotely Sensed Data, scheduled for June 23-28 at the Purdue University Laboratory for Applications of Remote Sensing, in West Lafayette, Indiana. Special emphasis will be on forest, range, and wetland assessment.

Papers are solicited on but not restricted to data correction and enhancement, digital classification techniques, evaluation of classification results, forest resources inventory, range/land assessment, wetlands and water resources, land use and geographic applications, crop inventory, soil survey, geology applications, georeferenced information systems, and technology transfer.

Four copies of paper summaries are due February 13. Summaries should be between 500 and 1000 words. Authors will be notified of paper selection by March 18. A camera-ready copy of the manuscript will be due June 26.

Send summaries to Douglas B. Morrison, Purdue University/LARS, 1220 Potter Drive, West Lafayette, Indiana 47906. \$8

**Magnetospheric Data Assessment**

The Scientific Committee for Solar Terrestrial Physics (SCSTEP) will sponsor a symposium at the Goddard Space Flight Center in Greenbelt, Maryland, May 21-23, to assess data gathered under the International Magnetospheric Study (IMS).

Objectives of the conference, which is scheduled immediately before the AGU Spring Meeting in Baltimore, are to determine the availability of data needed for correlative IMS studies; to examine the scientific results of the various workshops; to analyze the strengths and weaknesses of the various workshop forms; and to identify the work required to fulfill overall IMS objectives.

The first day of the symposium will examine the data available for correlative purposes. Of greatest interest are data sets that cover a large fraction of the IMS time period (1977 to 1979). The second day will concentrate on the examination of scientific results from previous IMS workshops. On the final day, the scientific progress of the IMS analysis will be examined.

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Send summaries to Douglas B. Morrison, Purdue University/LARS, 1220 Potter Drive, West Lafayette, Indiana 47906. \$8

## AGU

### Sponsors of 1980 New Members

Eight hundred fifty-five new members were elected in 1980. The following AGU members sponsored two or more new members last year.

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Meadows, Guy A.  
Vanick, Petr

#### Seven Members

Choviz, Bernard  
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Vogel, Thomas A.

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McLaughlin, Keith  
Nowlin, Worth D.  
O'Brien, James J.  
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Matumoto, Tosiatsu  
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Nur, Amos  
Rochester, M. G.  
Smithson, Scott B.  
Swanberg, Chandler A.  
Thompson, David E.

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Allen, John M.  
Allen, John S.  
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Anderson Jr., Donald E.  
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Bleher, Shawn  
Billington, Selena  
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Boville, Lawrence W.  
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Brown, Philip E.  
Bruhn, Ronald L.  
Butler, Robert F.  
Carstens, John C.  
Chase, Richard L.  
Chen, Chen-Tung  
Cheng, Ralph T.  
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Erickson, David M.  
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Fahnestock, David A.  
Farley, Donald  
Finnerly, Anthony A.  
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Frey, Fred A.  
Frohlich, Cliff  
Gang, Anthony F.  
Garland, G. D.  
Garmany, Jan  
Gelhar, Lynn W.  
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Gorman, John T.  
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Meyer, Robert P.  
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Morel-Soytous, Hubert J.  
Morgan, Paul  
Morgan, W. Jason  
Mori, Hirokazu

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Vogel, Andreas  
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Wang, Herbert  
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Whipple, Eiden  
Witanga, Peter J.  
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Wu, Francis  
Wyllie, Peter J.  
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Devries, Richard N.  
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Duce, Robert A.  
Dunning, Jeremy  
Eastman, Timothy E.  
Elthorn, Olav  
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Erickson, David E.  
Erickson, David M.  
Emat, W. G.  
Fahnestock, David A.  
Farley, Donald  
Finnerly, Anthony A.  
Forslund, David W.<

of Environment. (A. El-Shaarawi, Aquatic Physics and Systems Division, NWRI, Canada Centre for Inland Waters, P.O. Box 5050, Burlington, L7R 4A6, Ontario, Canada.)  
Oct 26-31 International Symposium on Quaternary Land-Sea Migration Bridges and Human Occupation of Submerged Coastlines, La Jolla, Calif. Sponsor, Center for Coastal Studies of Scripps Institution of Oceanography. (D. L. Inman, Director, Center for Coastal Studies, Scripps Institution of Oceanography, La Jolla, CA 92093.)

## 1983

July 18-23 Fourth International Conference on Permafrost, Fairbanks, Alaska. Sponsors, National Academy of Sciences, State of Alaska. (L. De Goes, Polar Research Board, National Academy of Sciences, 2101 Constitution Ave., N.W., Washington, DC 20418.)  
Sept. 12-14 National Water Well Association 35th Annual Convention and Exposition, St. Louis, Mo. (NWWA, 500 West Wilson Bridge Rd., Worthington, OH 43085.)

## 1981 AGU Spring Meeting

Baltimore, the site of the AGU Spring Meeting, May 25-29, is enjoying a major urban renaissance. Nowhere is this more apparent than in Metro Center, the 1000-acre downtown core of Baltimore. The convention center, an ultramodern meeting facility, is only a short walk from Harbor Place. Harbor Place is a skylight, terraced conglomeration of more than 20 water-side restaurants and over 100 boutiques.

**Hotel Accommodation.** A block of rooms is being held at three nearby hotels: the Baltimore Hilton, the Lord

Baltimore, and the Holiday Inn-Downtown. The Lord Baltimore and the Hilton are connected by a covered walkway. Read the housing application and MAIL THE COMPLETED APPLICATION FORM TO THE HOUSING BUREAU early to insure confirmation of preferred hotel.

**Registration.** Everyone who attends the meeting must register. Preregistration (received by May 8) saves you time and money, and the fee will be refunded if AGU receives written notice of inability to attend by May 15. Registration rates are as follows:

	Preregistration	At Meeting (after 5/8)
Member	\$45	\$60
Student Member	\$25	\$40
Nonmember	\$65	\$85

Registration for 1 day only is available at one half the above rates. Members of the American Meteorological Society, the American Society of Photogrammetry, and the American Congress on Surveying and Mapping may register for the meeting at the AGU member rates.

Students who are not AGU members should send in an application form with their registration payment. The difference between member (or student member) registration and nonmember registration may be applied to AGU dues if a completed membership application is received at AGU by August 3, 1981. Current AGU annual membership rates are: \$20 members; \$7 student members.

To preregister, fill out the registration form, and return it with your payment to the AGU Office. When payment is made by an organization, please attach the form wherever

possible; or be certain that your name and other pertinent information is on the check. Your receipt will be included with your preregistration material at the meeting. Preregistrants should pick up their registration material at the preregistration desk at the Convention Center. (On Sunday, from 5-8 P.M. in the lobby of the Hilton hotel.)  
The program and meeting abstracts will appear in the April 28 issue of *Eos*, which is mailed to all members in advance of the meeting.  
Complimentary badges for guests not attending the scientific sessions will be available at the registration desk.

## Social Events

An array of evening activities includes the Ice Breaker on Monday; the awards presentation honoring fellow scientists at a ceremony open to all participants, followed by a reception, on Tuesday; and an evening of fun and exploration on Thursday at the Maryland Science Center.

## Business Luncheons

There will be eight section luncheons: Geodesy, Geomagnetism and Paleomagnetism, Hydrology, Oceanography, Planetology, Seismology, Solar-Planetary Relationships, and Volcanology, Geochemistry and Petrology.

Check the appropriate spaces on the registration form and indicate number of reservations. Details of these activities will be published from time to time in *Eos*. Follow the Sail Inn Baltimore weekly update.

American Geophysical Union  
Spring 1981 Meeting

May 25-29, 1981  
Baltimore, Maryland

Mail this form to:  
Housing Bureau  
1 West Pratt St.  
Baltimore, MD 21201

## HOUSING APPLICATION FORM

## READ CAREFULLY:

Please print or type (pica spaced) all information abbreviating as necessary. Confirmation will be sent by the hotel to the individual named in Part I. If more than one room is required, this form may be photocopied.

## PART I

REQUESTOR	
LAST NAME	FIRST
NAME OF COMPANY OR FIRM	
STREET ADDRESS OR P.O. BOX NUMBER	
CITY	STATE
COUNTRY	ZIP-U.S.A.
AREA CODE	PHONE NUMBER

## PART II

**INSTRUCTIONS:** Select **THREE** Hotel/Motels of your choice from the list of participating facilities, then enter the appropriate code letters in the boxes below.

FIRST CHOICE	SECOND CHOICE	THIRD CHOICE
HOTEL CODE	HOTEL CODE	HOTEL CODE

**NOTE:** Rooms are assigned in "First Come First Serve" order and if none of your choices are available, another facility will be assigned based on a referral system arranged by your convention organizer. A cut-off date is in effect and your application may not be processed if received after 14 days prior to your arrival date.

\*AGU housing registration deadline is April 24, 1981

## PART III

**INSTRUCTIONS:** 1. Select type room desired with arrival and departure dates.  
2. **PRINT or TYPE** names of **ALL** persons occupying room.  
3. If more than two people share a room, check twin and the hotel will assign two double beds.

CHECK ONE <input type="checkbox"/> SINGLE (Room with one bed one person) <input type="checkbox"/> DOUBLE (Room with one bed two persons) <input type="checkbox"/> TWIN (Room with two beds two persons) <input type="checkbox"/> P+1 (Parlor plus one-bedroom suite) <input type="checkbox"/> P+2 (Parlor plus two-bedroom suite) <input type="checkbox"/> EXTRA PERSON	Arrival Date	MO DAY	Guest Names (Print Last Name First) 1. 2. 3. 4.
	Departure Date	MO DAY	
	Arrival Time	AM PM	

**IMPORTANT NOTE:** Hotel MAY require a deposit or some other form of guaranteed arrival. If so, instructions will be on your confirmation form.

SAIL INTO  
Baltimore  
AGU Spring Meeting  
May 25-29

## HOTEL ACCOMMODATIONS

PARTICIPATING HOTELS	HOTEL CODE	ROOM RATES
Baltimore Hilton Hotel 101 W. Fayette St. Baltimore, MD 21201 (301) 752-1100	BHOT	Single: \$43.00 Double: \$58.00 Twin: \$58.00
Holiday Inn-Downtown 301 W. Lombard St. Baltimore, MD 21201 (301) 685-3500	HIDT	Single: \$35.00 Double: \$38.00 Twin: \$44.00
Lord Baltimore Hotel Baltimore & Hanover Streets Baltimore, MD 21201 (301) 539-8400	LBDT	Single: \$33.00 Double: \$39.00 Twin: \$39.00
EXTRA PERSON	PARKING	
Hilton \$15.00	Hilton: nominal charge	
Lord Baltimore \$8.00	Lord Baltimore: nominal charge	
Holiday Inn \$7.00	Holiday Inn: free	
SUITES		
Hilton • Parlor plus one bedroom suite: \$125.00-\$190.00		
• Parlor plus two bedroom suite: \$250.00		

Be sure to enter the appropriate code letters on the attached form. Keep this sheet for your records, and forward the housing application form to the housing bureau at the address indicated.

All hotel reservations must be made on the housing form by April 24, 1981. No telephone requests will be accepted. Confirmations will be mailed directly to registrants by the individual hotels. After confirmation has been received, changes should be made with the hotel directly. Cancellations should be made with the housing bureau.

Any questions regarding your hotel accommodations should be directed to:

Housing Coordinator  
ABA Marketplace  
Baltimore Housing Bureau  
1 West Pratt St.  
Baltimore, MD 21201  
(301) 659-7000

PLEASE RETAIN THIS FORM FOR YOUR RECORDS

AMERICAN GEOPHYSICAL UNION  
1981 SPRING MEETING

Baltimore Convention Center  
Baltimore, Maryland  
May 25-29

## Registration Form

## PLEASE PRINT CLEARLY

Badge Identification

NAME ON BADGE

AFFILIATION

Name

Mailing Address

Telephone #

Address during the meeting if different than above

Please check appropriate box.  
Members of the cooperating societies may register at AGU member rates.  
Member badges are blue on white.  
Nonmember badges are red on white.  
☐ Member AGU ☐ Nonmember  
☐ Member cooperating society  
AMS-American Meteorological Society  
ASP-American Society of Photogrammetry  
ACSM-American Congress on Surveying and Mapping

Charge to: ☐ Payment enclosed  
☐ Visa ☐ MasterCard (Interbank)  
Card Number \_\_\_\_\_  
Interbank \_\_\_\_\_  
Expiration Date \_\_\_\_\_  
Signature \_\_\_\_\_

## RETURN THIS FORM WITH PAYMENT TO:

Meetings Registration  
American Geophysical Union  
2000 Florida Ave., N.W.  
Washington, D. C. 20009

Office Use  
Reference Number

DEADLINE FOR RECEIPT OF PREREGISTRATION—May 8, 1981  
Days you plan to attend: ☐ Monday ☐ Tuesday  
☐ Wednesday ☐ Thursday ☐ Friday

PREREGISTRATION (rates applicable only if received by May 8 deadline)

	More than one day	One Day
Member	<input type="checkbox"/> \$45	<input type="checkbox"/> \$22.50
Student Member	<input type="checkbox"/> \$25	<input type="checkbox"/> \$12.50
Nonmember	<input type="checkbox"/> \$65	<input type="checkbox"/> \$32.50

Abstracts April 28 (*Eos*) \$5

## SPECIAL EVENTS

Check the appropriate spaces and indicate number of reservations.

AGU AWARDS RECEPTION, following open presentation ceremony; includes food and drink, 7:30 p.m., Hilton  
Cost per ticket: \$9.25

SCIENCE CENTER: An evening of fun and exploration; includes food and beer, 8:30 p.m., Maryland Science Center  
Cost per ticket: \$6.50

## SECTION LUNCHEONS

Geodesy  
Geomagnetism and Paleomagnetism  
Hydrology  
Oceanography  
Planetology  
Seismology—cost per ticket: \$3.50 due to subsidy  
Solar-Planetary Relationship  
Volcanology, Geochemistry, and Petrology  
Cost of ALL LUNCHEONS, \$8.00 per ticket, unless otherwise noted

No. of tickets

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